

FINAL REPORT

Covering the Period

June 15 1948 to December 31 1953

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FIRST REPORT

Saving the Period June 15, 1948 to December 51, 1955

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Robert O'B. Sarpentur Ralph McDonough W. A. Patterson Norman Nappood, Jr.

Substitud in accordance with the terms of the contract noted shows.

Bruse H. Billings
Director of Research

Baird Associates, Inc. 35 University Read Combridge, Massachusetts

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SECTION I

Submitted under the Fermi of this Contract

1. INTEGRATION

This is the Final Report to be submitted under Contract NG-corr-60801, which was supported and supported at various and in various annuate by the Office of Naval Research, Bureau of Ordnance, Bureau of Ships, Watson Laboratories, AMC, Air Force Cambridge Research Center, and the Atumic Energy Gormission.

This contract was initiated and work started on 13 June 1988.

It is now scheduled to expire on the present expiration date of 31 December 1983.

The primary purpose of this contract has bosh to spensor research and development in the field of tunable narrow band optical filters. In its initial conception the aim was to develop and bring to practical realisation the tunable filter conceived by Dr. B. H. Billings, which uses narrow polarisation interference fringes arising from birefringent crystals to produce the narrow pass band (the Lyot fixed filter), with the tuning being added either by machinical rotation of some of the polarisar or crystal elements or by means of the electrically variable retardation which can be provided by the electro-optic effect in crystals.

During the first two years of the exertact almost all the emphasis was on the development of the fixed Lyot and tumble Billings

1. This conception was first described in a publication in the J. Opt. Sec. Am. 27, 758 (1947).

2. B. Lyot, Compt. rend. 197, 1593 (1955).

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filters to gractical realization. In the case of the fixed filter the development proceeded to the point where the apparetus was susceptible to practical manufacture. This filter differed from Lyot's primarily in the use of ADF in place of quarts and calcute. A manher of these fixed filters for the wavelength of Mx have proven very useful for studies of the son and its corone by astronomers. In the case of the tunable filter, although a number of research models were constructed, utilizing both mechanical and electrical tuning, its complexity and expense second an obstacle to its acceptance as a scientific tool for general way.

Towards the middle of 1950 there was somewhat of a shift in
the empireis of the work under the contract. Although continuing to work
on tumble merror band optical filters, the restriction to the polarization interference filter of the Lyot-Billians type was lifted. Various
modifications of the Fabry-Perot interference and interference in this
evaporated films, setal or dislectric, were considered and developed,
which might have considerable advantage in size, simplicity, and cost
over spectrometers and monochromators of equal resolution. Fundamentally,
this was the charactering of the original tumble filter. It was our
exived as being a simple and inexpensive replacement for a large grating
spectrograph, but it turned out to be about as costly and difficult to
build.

In addition the work was made more practical by applying the turnable flitter techniques to the scientifically and technically important field of isotope consentration measurement by spectrochamical techniques. Most of the practice was aimed at the hydrogen-deuterium problem as typical of the isotope measurement problem.

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The work under this contract has been reported in a series of Technical Reports which have been issued periodically. There are eight previous reports. It is felt that there is no edvantage to reprinting for this final report the detailed material covered in these reports. However, in the next section brief abstracts and suggestes of these reports will be given for reference purposes. Actually these interin technical reports were issued at appropriate times when it possed possible to treat a particular technical subject in a fairly complete may, rather than to issue them over equal working times. An attempt was made to make each one an entity in itself.

Accordingly, after these brief susseries, the bulk of the text of this final report will deal with the developments under this contract since the last Technical Report was issued on 1 July 1953, and with any other saterial not previously covered in the Technical Reports. The interested reader is referred to the earlier Technical Reports for detailed information as to the work on the early phases of this project.

2. AMETRACIS OF THE PRESENTAL REPORTS

All of these reports were issued under Scattreet Ed-cor-60801, with the Office of Maval Research as the principal educatebrative assempt.

2.1 2. E. Billings, "Summary of Tunable Filter Project", 30 December 1948.

This was a progress report which outlined the experimental progress then underway towards the development of fixed and tunable polarisation filters, under the following divisions:

2.1.1 Procurement and manufacture of elements for a rised Marrow Band Folkerisetics Filter.

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- 2.1.8 Devertigation of electrical tening methods.
- 2.1.3 Davelogment of mechanical tuning methods.
- 2.1.4 Improvement of the angular field of view of electrical tuning elements.
- 2.1.5 Development of fixed filters for isolating the spectral region to be covered by tuning.
 - 2.1.6 Operation in the infrared.
- 2.2 W. M. Danisin and S. J. Sage, "Problem of Fixed Filter Development for Tuning Purposes", il March 1949.

One of the unjor problems in the development of a treable filter is the manufacture of the besic fixed filter. A brief resum of the theory of the lipst type filter is described. The physical tolerances of the various elements is discussed, as are the techniques of orienting and eptically finishing the orystal plates, temperature control and lamination.

2.5 W. M. Braisin, S. J. Roge and J. S. Anderson, Rechnical Report dated 31 America 1949.

puried January 1 to August 1, 1949, in the development of a tenable server bend optical filter. The problem of mechanical processing of examents is discussed with more detail than in the last report, must of the successful techniques having been determined at this time. Sections on birefringence variation with temperature, temperature control and breating, wice corrector plates and their measurements; tuning alements and additional uses for filter components have been included. Finally, a section covering problems for future investigation closur this report.

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2.4 E. J. Weichel, Technical Datz, "The Index and Dispersion of Associate Dibydrogen Phosphete", El June 1950.

A compilation of the best values obtained from an evaluation of measurements at the Estional Eureeu of Stendards, University of Rochester; and Baird Associates, Inc. Final socuracy in index values is estimated as 5 in the sixth place beyond the decimal. Range 4000-6700 angetroms.

2.5 J. E. Billings, W. M. Braisin and S. J. Sage, Technical Expert dated June 1950.

This report covers the entire development of the fixed and tunable filters from the inception of the contract in June 1948 until June 1950. It thus reports or summerizes most of the information in the preceding reports on fabrication of the fixed filter. In addition, the development of the tunable filter is treated in detail for the first time. Several tuning methods are presented, some of which were tried in the laboratory.

2.6 3. N. Billings, S. J. Sags and N. Approof Jr., Technical Report dated 1 September 1972.

In this period the program was expended by extensive development work on other types of narrow band optical filters. The use of sultilayer dielectric reflectors to obtain greatly decreased band pass and increased transmission in the solid Fabry-Perot interference filter is described. Filters 50 augstrons wide with peak transmission of 80% were produced. The simple imagentains filters can be tuned over a range of at least 10% of the peak wavelength simply by tilting the filter. The

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two of these same multilayer reflectors to greatly improve the observateristics of the standard air-speed Pabry-Parct Interferences is discussed. Application was made to the extension of Fabry-Perot interferometry into the infrared and to the high resolution study of isotope abundances. The tolerances on interferometer plate flatness is discussed, and apparetus for flatness testing was devised and described in the report. Mice clearages silvered on both sides were developed as the best evaliable solution to the problem of obtaining thick, flat, solid spacer layers for Febry-Ferot filters with bend width the order of 1 engetrom. The difficulties of eilver reflecting layers in the infrared is discussed. The mice plates were then coated with multilayer reflectors designed for about 1.7 microns. Experiments on the 1.7 micron doublet of marqury demonstrated that high resolution Fabry-Perot interferometry could be extended to the infrared, with tuning by tilting. The explication of Fabry-Verot interferometry to isotope determinations is discussed, with emplication to the resolution of the marchy 5401 and the hydrogen 6560 isotope shifts and hyperfine structure. Finally, considerable further development on the timble polarization interference filter is presented, with emphasis on extending the tuning range of the filter to cover very wide epectral regions. This requires low order reterdation plates, and the available exterials of low linear or circular birefringence are surreyed and evaluated.

2.7 3. I. Billings, D. E. Bobinson, S. J. Sage and J. R. Welty, Toobsical Report dated December 1, 1952.

During the period covered by this report work centered on the application of the mice spacer Fabry-Perot interferenester to isotope analyzas.

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A possible instrument for analyzing hydrogen-deuterism mixtures is discussed. This instrument is based upon the fact that the mice spacer can be adjusted to produce two oppositely polarized light beams, one transmitting at Er and the other at Dr.

Experiments have been performed to test the sensitivity and reproducibility of the system. These preliminary tests indicate that the instrument could be used successfully to measure the relative concentration of two isotopes.

2.8 R. O'B. Carpenter, S. J. Sage and N. Repgood Jr., Technical Report dated July 1, 1955.

The Febry-Perot interferometer with mice cleavages as the spacer, as described in the last report, is completely investigated in this report. The properties of mice seem to render it impractical for the usage envisaged, namely, the variability from sample to sample in birefringence, the impossibility of cleaving to a prescribed thickness, and its absorption.

An ADP birefringent polarization filter, a unique type of tunable filter, is described and analysed completely in this report. It is found that it is suitable for only moderate accuracy isotopic analyses. Part of its difficulty arises from the fine structure and Doppler broadening of Mx and Dx, and this effect is analysed. The other difficulties are the high tolerances required in flatness of plates and in temperature control.

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3. PUBLISHED PAPERS AND TALES

The following published papers and talks presented at scientific society meetings were supported by or stimulated by the work of this contract, at least in part.

3.1 B. M. Rillings, "Tunable Merrow Passband Optical Filters", Winter Meeting of the Optical Society of America, February 20-22, 1947.
J. Opt. Soc. Am. 37, 561A (1947).

A polarization filter will be described which passes a region of the spectrum of the order of an angetrom in width and whose passband can be easily shifted to any wavelength in the visible spectrum. Previously, polarisation filters have been designed which have a linear aperture of several centimeters and a transmission benu an angetron wide. The action of these filters depends on the superposition of the polarised channel spectra formed by X cut plates of quarts between parallel rolarizers. Quartz plates were used whose thicknesses were in the ratios 1. 1/2, 1/4, etc., the alternate passbands of each plate were cancelled out by the black bands of the succeeding plate. Since the paraband of such filters was fixed in wavelength, they had limited application. Interest in filters of this type was restricted sainly to astronomers. Analysis shows that the 1, 1/2, 1/4 retardation ratios in the fixed Cl' ber can be wast considerably without inducing any appreciable change in its optical characteristics. Tuning is accomplished by eltering the rateriation of the successive filter elements so that transmission maxima in the individual channel spectra coincide at the desired passband wavelength. This retardation change can be made mechanically for example by stretching supplemental

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plastic sheets in series with the elements, or can be made electrically by using New cells or crystals with high electro-optic coefficients such as amonium dihydrogen phosphate. The additional reterdation acres has to oxceed a full wave. In the electrical case, a network sould be arranged so that with a single control the passband could be shifted through the spectrum. A tunable filter will be demonstrated.

5.2 B. H. Billings, "A Tunable Harrow-Band Option! Filter", J. Cpt. Ses.

Polarisation interference filters are described which can have a pass band renging from a fraction of an angetron to several hundred angstrons in width. The pass band can be shifted to any desired region of the spectrum. These tunable filters are tuned on the fixed filters discussed by Lyot and Evans. The transglusion band is formed by the superposition of the polarized channel spectra, produced by neout plates of quarts or other birefringent media placed between parallel polarisers. The birefringent plates have thicknesses in the ratio list 4 etc. The tuning to accomplished by changing the retardation of successive elements so that transmission maximu in the various channel spectra coincide at the desired wavelength. The reperdation change can be made machinely, for emergie, by stretching sum leasetal plastic shorts in series with the filter elements, or can be made electrically by using Kerr cells or crystals with high electro-cotic coefficients, such as essentian dibydrogen phosphate. The additional retardation never has to exceed a full ways at the wavelength of peak transmission. The assumed transmission of an experimental filter is shown. The electrical tuning method is particularly adapted to eathods

ray oscillograph presentation of spectra. The filter also has possible application in color reproduction and colorizatry. With a pass band of a half-angstrom line of sight motion of solar proximences could be determined by the use of the Doppler shift of the proximence radiation.

3.5 to 5.8 B. R. Billings and R. O'B. Garpenter. "The Electro-Ortic Effect in Universal Crystals of the Type X H.PO." J. Opt. Soc. Am.:

I. Theoretical, 59, 797 (1949)

When voltage is put across a crystal of the type X in the index allipsoid in the crystal is rotated and the length of the same changes. Two electro-optic coefficients are necessary to describe the effect in the most general plate of the crystal. The necessity for two coefficients is shown from consideration of the crystal symmetry. The behavior of Z-cut and X-cut plates is treated in detail. For X-cut plates with the voltage in the Z direction the crystal became biarial with the plane of the optic axes at 45° to the X and Y crystallographic axis. The retardation along the normal to the plate is directly proportional to the voltage and independent of the thickness. The characteristics of the polarization interference pattern for excited X and 2 plates are discussed.

II. Experimental, 59, 802 (1949)

In this paper it is shown experimentally that when a voltage is applied to the Z direction to a basel section of associate dihydrogen phosphate the crystal because biscial. The normal to the plate is the bisector of the axes and the plane of the axes is inclined at 45° to the crystallographic axes. The location of the axial plane is independent of the magnitude of the voltage. The retardation along the normal is directly

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proportical to the voltage and is independent of the thickness of the basal section. Shotographs are shown of the interference pattern obtained with polarized light when voltage is applied to the crystal.

The behavior of the crystal plate with an elternating voltage is also described. Between polarisers the crystal can operate as a light valve both for short pulses as in a centra shutter or for complex sound waves as in a sound-crefilm device. It is shown that with circular polarised high. To percent sinusoidal modulation introduces only 5 percent paraonic distortion.

III. Measurement of Soufficients, 40, 225 (1950)

The electro-optic coefficients of two crystals have been minerated by both static and dynamic methods. For ME_RPO_k, $r_{63} = 2.5$ k, $r_{k1} = 6.85$; for ME_RO_k, $r_{63} = 3.15$, $r_{k1} = 2.5$ 8 (units 10^{-7} c.g.s., 20° 8 k5560). The variation of the electro-optic response with the frequency of the applied field is given and related to the piezoelectric vibrations of the crystal. A theoretical relation for the difference in the electro-optic constants magnifed at constant stress and at constant strain is derived, which provides a check on the constant of the presently available electro-optic, where a check on the constant is and clastic coefficients for those crystals. The use of e.c. recibrds to obtain high precision in polarisation measurements is discussed.

IV. Angular Field of the Electro-Optic Shutter, 48, 18 (1958)

When voltage is emplied to a basel section of uniaxial orystel of the type II_FO, the crystal becomes biaxial. Between crossed polarisers it can be considered, then, to not as a light valve. The angular field

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of view is limited by the natural retardation of the crystal. This paper discusses without by which this natural retardation can be minimised or effectively canceled so that the retardation with no voltage is essentially some over a large angular field. In the first method a basal section of uniaxial crystal of opposite sign is placed in series with the slab to be excited. Another technique uses two slabs of similar crystal with a 50° optical rotator placed between them. Both these techniques are treated theoretically and experimental resourcements are given to show the resulting angular polarisation pattern. This pattern is shown both for the electrically excited case and the uncroited case.

5.7 R. 0'B. Carpenter, "Induced Birefringence of AIP and KT/ as a Function of Temperature", Maeting of the Optical Society of America, Cleveland, October 26-88, 1950; J. Opt. Sec. .m. 40, 7984 (1950).

In strongly paraelectric and ferroelectric crystals the photoelastic and electro-crytic coefficients, along with the elastic, dielectric
and photoelectri; properties, ambiet large and enouslous temperature
dependence. Data have been obtained on the variation with temperature
of the electro-crytic r_{60} and the photoelestic q_{60} coefficients of AEP and
EEP from room temperature down to their respective transition temperatures,
i.e., in the para-electric region. Several techniques of making these
measurements will be described and compared as to sommany. The elastical
thermodynamic theory of the interaction of electrical and mechanical

^{5.} W. P. Mason, Phys. Rev. 69, 175 (1946).

^{4.} Dricker and Scherrer, Belv. Phys. Acta 17, 346 (1944).

^{5.} Surpenter, J. Opt. Soc. Am. 10, 225 (1950).

influences within the crystal will be applied to this data in an attempt to relate the induced birefringence to a properly chosen set of meansscopic parameters in a temperature-independent fushion.

3.8 B. H. Hillings, S. J. Sage and W. M. Arrigin, "A One-American Fasse-Rend Polarisation Interference Filter", Winter Meeting of the Optical Scatter of America, New York, March 9-11, 1930; J. Opt. Soc. 28. 40, 2024 (1930).

One of the main difficulties in the construction of a narrow bend Lyot-Remn⁶ type polarization interference filter is the procurement of large slabe of transparent birefringent meterial. Previous filters have been built mainly of quarts. The narrower the bend width, the thicker the required birefringent alabe. A lA width would need a quarts thickness of approximately 8.6 inches. Optical quarts of this thickness is essentially unavailable in quantity. A filter has been constructed using emonion dihydrogen phosphate as the birefringent material. Although this crystal is easily available it is extremely difficult to work because of its softness and because it is water soluble. Furthermore, it has to be held to finer tolerances then the quarts slabs because of its higher birefringence. The filter which will be shown has mechanical tuning which emables it to be adjusted over a range of about \$A\$. It is planted in future work to type 7 these filters both electrically and mechanically over a wide spectral rusion.

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^{6.} B. Lyot, Scapt. rend. 197,1595 (1955); Y. Chann, Mature 141, 291 (1958).

^{7.} B. E. Billings, J. Opt. Soc. &s. 37, 758-746 (1947)

3.9 3. N. Millings, "A Birefringent Frustrated Sotel Reflection Filter", J. Opt. Soc. Am. 40, 471 (1950).

In the standard Turner Frustrated Total Reflection filter the pass bands occur in pairs which are oppositely polarised. The maximum theoretical transmission in any band is thus 50%. In this paper it is shown that if the high index layer of the filter is unde of a birefringent material the oppositely polarised pass bands can be made to coincide so that the maximum theoretical transmission becomes 100% and the bands occur singly. Measurements are shown on a filter made with such a birefringent layer.

- 5.10 B. H. Billings, "Involuntary Polarisation, and Depolarisars", Invited
 Paper, Winter Newting of the Optical Society of America, Washington,
 March 1-3, 1931; J. Opt. Soc. Am. 41, 2007 (1931).
- 3.11 B. H. Billings, "Harrow Band Optical Interference Filters", Photographic Engineering 2, 45 (1951).

A publication of some of the principal results described in the same Catalled tasked on compart, reference \$40.

in the Infrared), J. page. or redict . . . 507 (1950).

Thin films have found use in infrared light as optical filters, as bolometers and for reducing the amount of reflected light. A thin metal film can also be designed so as to wholly absorb a given wavelength in the infrared bands.

Two types of filters are described. One is Dr. Turner's frustrated reflection filter, which transmits a narrow band in the wigible 些

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spectrum. The present author has extended its use to infrared and to ultra-short waves of 1 on wavelength. A sample is demonstrated, transmitting a hand 0.1 μ wide in the neighborhood of 5μ .

The second type is the Demnison Eadley reflecting filter, which reflects a set of infrared bands; several such filters are shown. The principle of the Desmison Eadley filter is the following: a thin metal film of 577 chms/square resistance absorbs all the energy in a given wavelength, when placed a quarter wavelength in from of a perfectly reflecting mirror.

Such a thin film can constitute a bolometer, thus obviating the difficulty of having a satisfactory black body (all metallic blacks are gray in far infrared light): Such a black is no longer resoled when the thin film is of appropriate resistance.

3.13 W. H. Dreisin, S. J. Sago, J. M. Anderson and J. C. Jelatis. "High Speed Photography with the Electro-Cytic Shutter", presented at 3 pasting of the Her England Section of the Optical Society of America, April 10, 1952.

The use of an artificially grown erretalline natural, assentian dibytrogen phosphate, as a high speed light value is described. The cariestation required to produce the optimum electro-cytic careet, types of electrodes used, electronic requirements of the pulser unit activating the crystal, and the general performance of the shutter were included. Photographs of exploding firecreakers, exposed by the light of the explosions themselves, with exposure times controlled by the shutter and varying from 50 to 10 microseconds were shown. An actual deconstruction

SESTICE II - A TRIBETCAL EMPORT

(This work was principally done in the period June 1 to December 31, 1973

The Fabry-Ferot interferenter is an entrantly simple dispersing instrument requiring only two flat quarts plates with reflecting layers on one side held together by some form of mount to insure parallelism of the plates. In spite of its simplicity, it has long been recognized as the instrument of the highest spectral resolution swallable for the spectroscopy of close line structures. In this report its application to the problem of the spectrochemical analysis of hydrogen-deuterium mixtures is surveyed in detail.

The separation of the plates is determined to be 0.24 inches by the requirement that the deuterium fringe be exactly sidney in the free spectral range between two successive hydrogen fringes. At such narrow spacings the instrument is not at all delicate or difficult to align, as compared to the more difficult problems arising in closer line structures, where the plates have to be spaced several inches apart.

In order to seen the wavelength spectrum, the interferences is mounted in a vacuum tank. The fringes are observed through a circular aparture and photocell placed at the center of the well-known Jabry-Perot circular fringe pattern. By varying the air pressure in the tank, the fringes are collegued, because of the varying index of refrection, and the wavelengths are scanned past the aparture. The resulting intensity varsus wavelength is plotted on a strip chart recording microsmater. For air the scanning is only about .57 orders (or free spectral ranges) from

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vacuum to one etmosphere. This was improved by using 60 (.55 orders) and exlipher homefluoride (1.5 orders).

The theory of the Pabry-Perot fringe shape is discussed. In any dispersing system the resolution and the contrast are the important parameters. The contrast is of considerable importance if deuterium is ordinary water (.016%) is to be detected, for the Sauterium fringe is superposed on the hydrogen background.

After determining the theoretical factors which influence the resolution and contrast, an extensive experimental progress was undertaken to achieve the maximum values of these two important parameters. The reflectivity of the reflecting layers, the flatness of the plates, and the width of the lines in the sources are the three limiting factors on resolution and contrast.

It is shown that multilayer reflecting layers are far superior to silver, and the multilayer prepared gave more than adequate reflectivity, resolution and contrast. The reflectivity was then less a limitation than the flatness of the best plates which could be obtained. A considerable discussion is given of the techniques of making multilayer reflectors, and of the apparatus which was constructed for the independent and precise measurement of the reflectivity.

It was necessary to considerably improve our apparatus for the independent determination of the flatness of the plates. Actually, in flatness determinations the some limits on reflectivity hold, so that it was necessary to first solve the reflectivity problem before the flatness of our excellent plates could be observed. When this was finally done,

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it was found that the plates were flat to somewhere between 1/50 and 1/100 of a fringe of red light.

At this point the width and contrast of our fringes were principally determined by the actual source width. A definite assymmetry in the H alpha line was observable, which is due to the Doppler broadened fine structure. The direction and magnitude of the asymmetry are in accordance with the structure, and the width is in accordance with a reasonable assumption as to the effective Doppler temperature in the source.

The recorder has to be assessed improved in order to achieve the maximum signal to noise ratio. This involved adding additional input gain or sensitivity, and increasing the time constant by adding capacity to the input. Some care had also to be exercised to reduce the stray room light to a minimum, since any fluctuations in the strays would be disturbing and the deuterium light at such low concentrations is very weak. It is or course only the fluctuations in the dark current and instrumental background which limit the sensitivity of the method.

By means of all these experimental efforts, it was possible to observe the deuterium fringe in ordinary water, an effective contrast of about 6000 to 1.

A section in the report is devoted to a comparison of the light gathering power, or luminosity, of the Febry-Perot with prises, gratings and other dispersing elements. The high luminosity of this simple instrument gives it a further advantage over the other instruments. This is of particular importance when the application is energy limited. In this

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connection the multilayer reflectors are vital, for although silver can be everywheld to very high reflectivity; this is achieved at considerable expense in light transmission at the peak.

The report also discussed the problem of obtaining sources of high stability and nervow lime widths. A mixing and discharging apparatus for hydrogen and deuterium gases was constructed. The most convenient source appears to be the veter vapor discharge, and the final results were obtained with this source.

1. IFTEODESTION

For some months work under this contract has been concerned with isotope analysis by means of methods closely analogous to standard spectrocharical techniques. In particular, a number of simple interference methods have been investigated in the hope of deviating instruments for separating closely spaced storic lines originating from different isotopes of the semi-element which would be less builty and expensive than large grating spectrographs.

In previous reports a polarization interference has been described which asparates two lines in a simple two isutope problem into two beams polarized at right angles to each other. This led to an ingenious method where the wavelength— were not dispursed in ordinary space, but only in polarization space. Calculations were made to show how the system might work for hydrogen-deuterium analysis, as a simple, yet

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important, example of the two isotope problem. A fundamental difficulty in applying the sectod to low concentrations or with high accuracy arose in any case where the two lines had structure, which might arise from the fine structure of the lines or from the line broadening in the source. This difficulty arises in principle from the fact that in a sense there are only two polarization states, because any other state of polarization can be resolved into two components in the given states (usually taken at right angles). Therefore the polarization method, without additional dispersion in ordinary space, can only resolve two purely monochromatic lines.

Some work was also done on a mice Fabry-Ferot interferometer system. This method was found to be impractical because of the great difficulty in selecting mice cleavage plates which had both birefringence value and thickness to the required tolerances.

The Fabry-Ferot interferometer produces a suries of very sharp interference fringes and produces a dispersing instrument of very high resolution and dispersion, but with a limited free spectral range because of overlapping orders. Therefore it is very highly suited to spectrum analysis problems involving only a very few very closely spaced spectrum lines, which is essentially the problem of isotopic analysis if all other lines in the spectrum can be eliminated with auxiliary filters.

Since mice is birefringent, i.e., has two indices of refraction, it produces two independent sets of Fabry-Perot fringes. Thus in the sense of the words need above, it is an instrument producing both dispersion and polarisation. Since the polarisation double bear method led

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to difficulties when there were more than two pure mozochromatic compoments, reconsideration of the simple air spaced Fabry-Peret interferemeter was undertaken.

Perot for spectrochemical analysis lies in the fact that two circular elits of different radii would be required to select two lines, and that precise tuning of the spacer thickness would have to be maintained in order to keep the two lines centered on the two slits. There is, however, a method which has been described in the literature for precise tuning of the fringes, which consists of varying the air presents within the spacer layer. Since this varies the index of refraction of the spacer layer, the circular fringes are collapsed and expended by this method. This led to the idea of a simple Fabry-Perot interferometer enclosed in a vacuum chamber, with the wavelengths passed by a single sparture controlled by varying the air pressure within the chamber.

In the next section this interference will be described in detail.

2. CHENNAL PROPRIETION OF THE PARTY-PERCY DEVELOPMENTS.

A Fabry-Ferrit interferometer consists of two highly reflecting layers separated by a spacing layer which maintains the two reflectors accurately place and parallel. When six is the spacer the two plates are hald spart by a three-legged device called a "spacer".

The theory and practice of the interferometer are described in detail in Tolensky's "High Mesolution Spectroscopy". Only the most 8. C. Defour, Thesis, Paris (1950).

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important equations of its behavior will be reproduced here for reference, and for considerations of the characteristics of the instrument which are critical for the H-D application.

Constructive interference occurs and bright sharp fringes are found at those wavelengths A and angles & (from the normal) satisfying

$$m\lambda = 2\mu + \cos \phi \tag{1}$$

where n is an integer, called the "order number" of the fringe, μ is the index of the spacer layer, t is its thickness.

A photograph of a typical set of circular fringes taken in momechromatic Hg 198 5462), light is shown in Figure 1.

Differentiating

$$\frac{dn}{d\lambda} = -\frac{dn}{\lambda^2} \cos \theta \tag{2}$$

A parameter of intervet is the so-called "free spectral range", the specing in wavelengths between successive fringes of the same wavelength, i.e., & corresponding to dn = 1.

Free spectral range
$$(\Delta \lambda)_2 = -\frac{\lambda^2}{8\mu t \cos \theta} \approx \frac{\lambda^2}{8t}$$
 (5)

The expression for the free spectral range in wave numbers is simpler, being independent of wave number

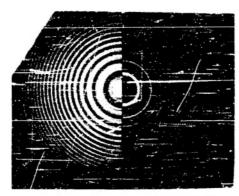
$$(\Delta^{\mu})_{f} = \frac{1}{2\mu t \cos \theta} \approx \frac{1}{2t} \tag{4}$$

Now consider the application of the interferometer to hydrogendeuterium analysis with E_{CP} where $E_{CP} = D_{CP} = 1.786$ angetroms. In order to



MIGURE '1

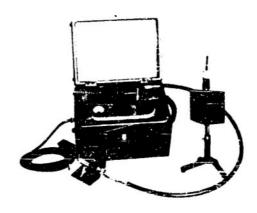
MERCURY 198 SOURCE AND MICROWAVE EXCITER



ORDINARY

MERCURY 198

FRINGES TAKEN WITH FABRY-PEROT INTERFEROMETER HAVING 43MM SPACER DESCRIBED IN TECHNICAL CIRC. 502-1.



LAMP MOUNT & MICROWAVE EXCITER

The Mercury 198 source and exciter consists of two separate assemblies:

- a. The Mercury 198 electrodeless lamp containing approximately 0.2 mg. of Mercury 198 which has been prepared by transmutation of gold in a pile and contains no other isotopes. Argon is used as a buffer to enhance the mercury excitation. The lamp is 6 mm. in diameter by 14 cm long and is made of Vycor #7911 glass.
- b. The Mercury 198 Microwave Exciter operates di a wavelength of 12.2 cm in the 2400-2500 megacycle band. Radio frequency energy is generated in a continuous-wave magnetron oscillator and is carried in a flexible coaxial cable to the director which concentrates the external radiation on the lamp. The unit has a maximum power output of 125 watts. Relative power output is indicated by a per cent of maximum power meter on the front panel.

High-voltage direct corrent for the magnetron is provided by an AC-operated power supply which includes separate plate and filament transformers, two 816 rectifier tubes and a filter circuit. The input to the high-voltage transformer is controlled by a time delay relay, a timer and a Variac which controls the RF power output to the magnetron. The from pane! Preheat and Power Switches control the application of filament and plate power respectively, while adjacent panel lights indicate switch positions. An AC operated motor-driven blower supplies cooling air throughout the cabinet.

The unit is portable and is packaged in a hand-rubbed quartered cak cabinet measuring 15" x 10" x 11". The exciter is designed for operation from a 60 cycle 115-volt AC power source.

The lamp mount has an aperture of $\frac{1}{2}$ " x $l\frac{1}{2}$ " and is furnished complete with pedestal and water jacket.

TECHNICAL CIRCULAR RD 502-3 Revised

Baird Associates, Inc.



PRECISION INSTRUMENTS FOR ANALYSIS AND CONTROL

33 UNIVERSITY ROAD, CAMBRIDGE 38, MASSACHUSETTS

obtain the best possible separation, it is desirable to have the two wave lengths separated by exactly one-half order.

$$\Delta n = n_{\underline{H}} - n_{\underline{D}} = \frac{1}{2} = 2\mu \ t \left(\frac{1}{\lambda_{\underline{H}}} - \frac{1}{\lambda_{\underline{D}}}\right)$$

This determines the required spacing t between the two plates.

$$t = \frac{1}{k\mu \, \Delta \epsilon} \tag{5}$$

where $\Delta \delta$ is the difference in wave number for the two lines. For $\frac{\pi}{\alpha} = D_{\alpha'}$ $\Delta \delta = 4.145 \text{ cm}^{-1}$ and

$$t = .0605 \text{ cm} = .0237 \text{ inches}$$
 (6)

There are of course other solutions separating the two fringes by 3/2, 5/2, etc. orders, and these require thickness 5, 5, etc. times as thick.

Econver the tolerances on alignment, spacing, temperature control, etc., decrease with thickness, so the chosen solution is considered the simplest in practice.

Now the question arises as to how many fringes does the pattern shift from atmospheric pressure to vacuum. The change in order number is

$$n_{elr} - n_{vac} = \frac{2(\mu - 1)t}{\lambda} \tag{7}$$

For air, $\mu - 1 = .00029$, so

Thus if the spacer thickness is adjusted to center the M-frings exactly

in the slit at atmospheric pressure, the D-frings can be tuned through the slit by evacuating the chasher.

The maximum value of cos 0 and of the order number n occur at the center of the circular fringe pattern. For t=.0605 cm, n_0 is about 1857 tuves. Integral orders from the normal cooper at about 1^049° , 2^041° , 5^015° , etc.

The siventege of the pressure tuning method, of course, is that one may pick up the light photoelectrically from a circular sporture at the center of the pettern, and by varying the pressure in the instrument scan the spectral vavolengths past this aparture, recording the transmitted intensity versus pressure if desired.

Accordingly, a vacuum tank for holding the interferometer was designed and constructed. It was discovered that there was a scrap brase cylinder in the netal stores which case so close to the desired dimensions that a considerable saving in construction expense could be achieved by starting with this cylinder. The drawing for reworking this cylinder is shown in Schematic ER-2126. Figure 2 is a photograph of the finished tank. If it were not for using the available brase cylinder, the dismeter could have been made acceptant smaller.

The simple ameroid type pressure gauge is an Ashcroft gauge manufactured by Hamming, Harvell and Moore. It presents atmospheric pressure to vacuum on a scale of 30 inches. The other fittings were an inlet for the vacuum pump and a small needle valve for admitting air. A simple Welch pump is adequate for this instrument. Holes were drilled in both the end plates and collimating leases were scaled in with Apieson "Q"



The state of the s

putty to serve as both window and lens. These are not shown in the figure, but were added later. The input condenser lens was of about finehes focal length, and is not required to be of particularly high quality. The focusaing lens was a fine achrometic doublet of 100 cm focal length.

The interformmeter is the standard Baird Associates, Inc.

Febry-Perot Interferenseter. This instrument was originally designed similar to principles suggested by Mr. F. O. Westfall of the Baticual Bureau of Standards. A photograph of the Baird instrument is shown in Figure 3.

The specified etalon specing is .0237 inches. This is much conclient them any of the standard invar spacers furnished with the instrument. It was pointed out in Tolansky's book that in the manufacture of ordinary steel balls for bearings, the diameters are held to such close tolarances that these serve as the bost practical spacers for Tabry-Perot work at very low separation. Our spacing was also somewhat smaller than is usually found in steel balls, but inquiries to the various steel ball manufacturers uncarthed two balls close to the desired diameters. 1000 of each size were ordered: .0312 inches (.78 mm) and .0256 inches (.65 mm).

A spacer plate had to be made to hold the balls from rolling around. The first one was made of .020 brass and is shown in Figure 4.

Trouble was had with buckling, and better results were obtained by cutting a similar piece out of a mica classage.

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FIGURE 3

APPARATUS FOR HIGH DISPERSION INTERFEROMETRY

FABRY -- PEROT INTERFEROMETER

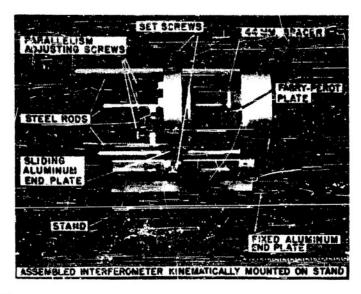


Figure 1

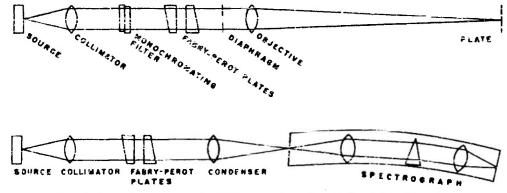
General.

The Fabry-Perot interferometer is one of the simplest high resolving power instruments which are available to the research worker. When the plates are coated with suitable materials the instrument is useful in the infra red, visible, and ultraviolet.

The Fabry-Perot interferometer has many uses: the determination of absolute wavelengths, the refractive indices of gases; the measurement of lengths, the widths of spectral lines; the evaluation of hyperfine structure.

The theory of the Fabry-Perot interferometer is dealt with at length in the references given in the bibliography. Two illustrations of how it fits into an optical system are given below.

Features of the Baird Associates Fabry-Perot Interferometer are the remarkable east of assembly, a minimum of adjustments, and a holder which accepts spacers from 0 to 100mm.



FABRY-PEROT INTERFEROMETER CROSSED WITH SPECTROGRAPH

Baird Associates, Inc.



PRECISION INSTRUMENTS FOR ANALYSIS AND CONTROL

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Description of Instrument.

The Baird Associates Phbry-Perot Interferometer consists of two fused quartz discs with flat surfaces, spacers to hold the plates accurately parallel a fixed distance apart, and a holder to keep the plates lightly pressed against the spacer ends and to hold this assembly in an optical system.

The plates, Fig. 2F, are 21 in diameter and .625" thick. They are wedged 5 to 10 minutes of arc in order to displace the weak secondary pattern formed by the uncoated faces. One face of each of the two plates is figured to at least $\lambda/40$ measured at the mercury green line within a 40mm diameter circle concentric with the bevelled edge. The opposite side is figured to 1 wave. An arrow placed at the thickest part points toward the more highly figured face. The plates are supplied uncoated. For use a highly reflective, partially transmitting coating is deposited on the figured face. The coating is kept within a 34mm diameter circle so that the spacer ends rest on clear plate. The best performance is obtained with a silver coating. Baird Associates can supply silver-coated Fabry-Perot plates and will recoat plates on request (See price list). Silver coatings often have deteriorated substantially several weeks after deposition by the formation of sulfide. Because of this many users will wish to have coating facilities available locally.

The spacers, Fig. 2 C, D, and E, which hold the plates parallel a fixed distance apart are made of three invar rods press fitted into one or two stainless steel rings 2.75" in diameter. The spacer ends are optically finished to 1/10 of a wave of mercury green light over an area at least 1 square millimeter.

When the polished spacer ends are in optical contact with the inner surfaces of the Fabry-Perot plates a yellow interference color is seen that permits easy checking of assembly. Standard spacer lengths are 7mm, 17mm, and 44mm. Other lengths up to 100mm are available on special order. The variation in spacer leg length is less than $3\lambda/2$ of mercury green light.

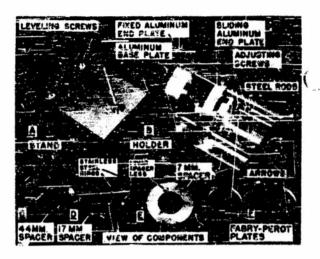


Figure 2

The holder, Fig. 2B, which keeps the two plates lightly in contact with the ends of the spacer, consists of an aluminum base plate on which is mounted a fixed aluminum plate. Into this fixed end plate are press fitted three 1/2"steel rods. A second aluminum end plate slides on the three steel rods. One Fabry-Perot plate rests against 3 pins in a depression in the fixed end plate. The clear aperture of this end plate Is a hole 1.375" in diameter. The other Fabry-Perot plate fits into the sliding end plate and is held in by three springs. The spacer fits inside the steel rods. Set screws clamp the sliding end plate in position and screws adjust the spring tension on the three spacer legs for the final parallelism adjustment.

The holder has a cone-slot-plane base (Kinematic mounting) for the precise repositioning on the ball bearing levelling screws of the stand. The outside dimensions of the interferometer are $6.75'' \times 4.875'' \times 4.375''$. The interferometer; stand, pair of Fabry-Perot plates; and 7mm, i7mm, and 44mm spacers fit into an attractive, compartmented oak case. Outside dimensions of case are 12.5" x 11" x 6.25". See photographs.

References

- 1. Tolansky, S. "High Resolution Spectroscopy" Menthuen and Co. Ltd. London 1947
- 2. Candler, C. "Modern Interferometers" Hilger and Watts Ltd. Hilger Division 1951
- 3. Meissner, Karl Wilh. Journal of the Optical Society of America Volume 31 (June 1941) pp 405-427 "Interference Spectroscopy Part I"
- 4. Williams, W. Ewart "Applications of Interferometry" E. P. Dutton and Company Inc. c. 1930



The optical arrangement is shown in Figure 5. The fringer are in focus at the focal point of the 100 cm collimator. At this point a demon variable tris disphrage was placed. A field loss and field stop gathered the light into the 1921 photosultiplier. These served to eliminate all stray room light not coming from the immediate vicinity of the collimator loss. The photosultiplier power supply could supply various valtages stepwise. However after contracts of several hundred were obtained, it became necessary to build a room light shield around the instrument, which was not in a dark room. A very entisfactory Leeds and Borthrup Speedomax was available at the laboratories for recording the light as the pressure was varied in the tank. This instrument has standard scales of 2, 5, 10, 20, 50 and 100 microsuperes full scale. A Resson microsuperer with 10 and 100 microsupere scale was also occasionally used.

A refinement which has not been added to the apparatus as yet is a synchronization between the pressure in the tank and the recorder chart paper drive. This was done by Jacquinot and Dufour in a similar apparatus.

5. VALUE INTEGRATE DISTRIBUTION - CONTRAST, INSCLUTION AND TRANSMINION OF THE PRINCES

Nolaneky gives the general empression as first derived by Airy for the friege intensity distribution.

$$I = \frac{T^2}{(1-R)^2} \frac{1}{1 + \frac{4R}{(1-R)^2} \sin^2 \frac{\delta}{2}}$$
 (8)

9. F. Jacquinct and C. Dufour, J. Recharches du S.N.R.S., Bellavue, 6, 1 (1948)

SCHEMATIC OF OPTICAL ARRANGEMENT FABRY-PEROT INTERFEROMETER WITH PRESSURE TUNING

FIGURE 5

where $\delta = (2x/\lambda)$ & t cos 0.

A plot of this equation is shown in Figure 8, with the quantities of interest merbed.

The ratio of the maximus transmission to the minimum is defined as the "contrast" of the fringes.

$$e = \frac{T_{\text{max}}}{T_{\text{min}}} = 1 + \frac{\frac{2\pi}{(1-R)^2}}{(1-R)^2} = \frac{(1+R)^2}{(1-R)^2}$$
 (9)

This factor is important in determining the concentration accuracy of the instrument, or its constitutity in measuring very less concentrations of deuterium, since the D light will be lost in the background of the M fringe. The minimum intensity occurs balfvay between maxima, which is one of the principal reasons for choosing an etalon spacing making D appear one-half order from M.

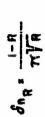
Representative values of the contrest are:

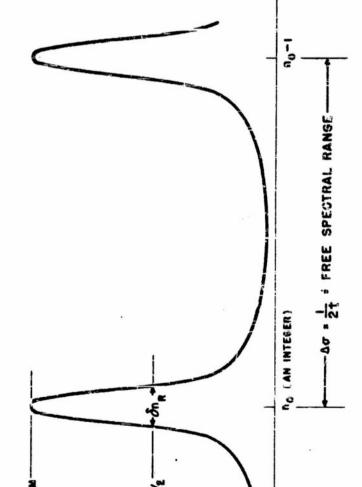
3	Contrast	Inip/Inex
70%	31.0	3.225
75%	49.0	8.00%
80%	81.3	1.23%
85%	152	0.66%
90%	36 0	0.29%
99%	1650	o.06≸

It should be exphasized that the contrast depends only on the reflectivity R and not on the transmission or absorption in the reflecting layer.

C

$$T_M = \left(\frac{T}{1-R}\right)^R$$





FRINGE INTENSITY DISTRIBUTION

FIGURE 6

(

When we come to discuss the reduction of electronic noise the effect of the contrast on the resolving power in the intensity scale will be considered more carefully. It is the fluctuations in the background, not the background itself, that limits the accuracy and sensitivity in intensity measurement. In spectrum analysis one plots light intensity versus wavelength. The contrast has an important bearing on the resolution in the intensity scale. We now consider the resolution in the wavelength scale, which we will call by the limited term "resolution".

The wavelength resolution is limited by the half width of fringes which can be obtained. There have been a number of criteria or conventions used for the resolving limit of dispersing instruments, but they are all about the same. It seems most convenient to take as criterian that two lines can be just resolved if they are separated by total width between half maxime for either. This criterion is not greatly different from the well-known Rayleigh criterion.

From equation (6) for the fringe intensity distribution it is easy to show that the intensity is one-half when the order number n = 24 t cos o/h departs from an integer by

$$(8n)_{1/2} = \frac{1-R}{2\pi\sqrt{R}}$$

The total width between half maxime, measured in fractions of an order of interference, is just twice this, and will be designated by

$$\delta n = \frac{1-R}{s \sqrt{R}} \tag{10}$$

1

This quantity is the resolving limit in orders of interference, and is a quantity depending only on reflectivity, not on wavelength, etalon specing, nor angle of observation. Jacquinot and other French workers in the field have found it convenient to define the "fineses" I as the reciprocal of this resolving limit. This is a convenient parameter to be taken as the figure of merit of a Fabry-Perot interference.

$$\bar{\pi} = \frac{1}{8n} = \frac{\pi \sqrt{R}}{1 - R} \tag{11}$$

The resolving limit in wavelength or wave number is now readily obtained by multiplying on by the free spectral range in wavelength or more number.

$$\delta\lambda = \delta\alpha \Delta\lambda = \frac{\Delta\lambda}{B} = \frac{\lambda}{B_0} = \frac{\lambda}{B_0} \approx \frac{\lambda^2}{2t B}$$

$$\delta f = \delta \alpha \Delta f = \frac{\Delta f}{B} \approx \frac{1}{2t B} = \frac{1}{2t} \delta n = \frac{f}{B_0}$$
(12)

where n = the order number.

The "resolution" as ordinarily defined in optical problems is now

Resolution =
$$\frac{\lambda}{5\lambda} = \frac{5}{50} = \frac{20}{50}$$
 (13)

The "finesee" is also a number of fundamental interest in the behavior of a Fabry-Perot interferencer, because it is approximately equal to the number of interfering beaus of approximately equal intensity in the multiple beam interference. In other words, the beaus of number greater than H have been diminished by successive reflection until they make a negligible contribution to the sum of interfering amplitudes. In

the Lammer-Gehrke interferometer, for example, there are only a finite number of beams, but all have unity amplitude, since they are reflected at total internal reflection. A Lummar-Gehrke plate arranged to produce interference between just N beams has similar properties to a Fabry-Perot interference with finesse N.

The quantities now which depend solely on the reflectivity are the contrast 0, the resolving limit on, the finesse N. These are plotted versus reflectivity in Figure 7.

In the case where R is .90 plus or minus .10, there is an approximate relation between 6 and R.

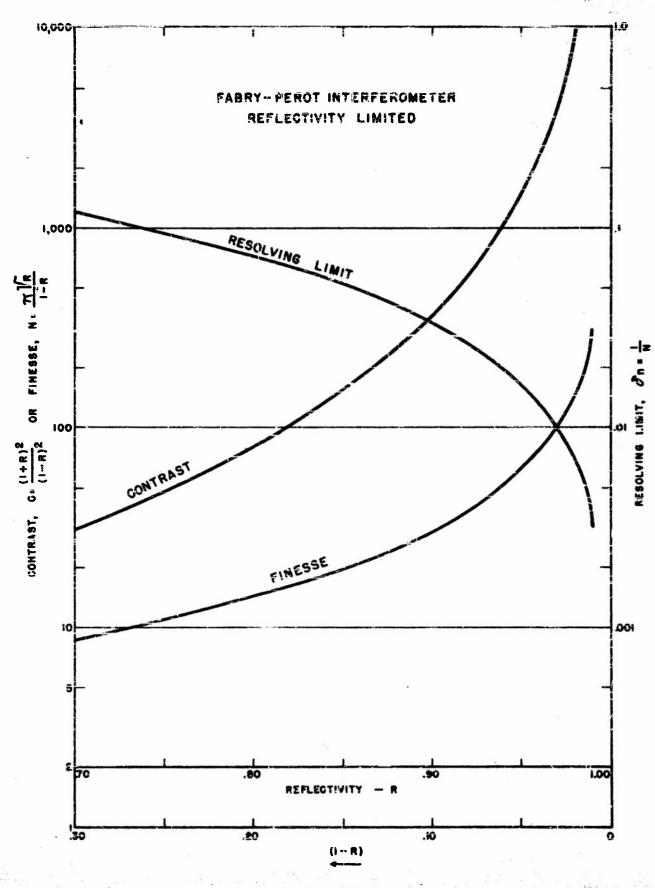
4. THERTICUES OF MAKING REPLECTING LAYERS

The relative shape of the interference fringes in the Fabry-Ferot depends only on the reflectivity of the layers. If there is absorption in the reflecting layer, as is always the case with thin metallic films, the transmission of the fringes is reduced everywhere by the same factor, the first term in equation (8). Thus for the peak transmission of the fringe, we can write

$$T_{M} = \frac{T^{2}}{(1-R)^{2}} = 1 - \frac{2A}{1-R} + \frac{A^{2}}{(1-R)^{2}}$$
 (14)

since the abscrption A = 1 - R - T.

This loss in transmission can of course become serious if one is operating in an energy limited situation. Such is very likely to be the case of one attempts to observe the deuterium alpha line from a



Secretarian Company of the Company o

FIGURE ,7

R

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sample with only a few hundredths percent deuterium. For the hydrogen discharge will be week to start with if its line broadening is to be kept small, and photomultipliers have considerably reduced sensitivity at \$500 engaintees.

In Figure 8 is plotted the transmission T_m of the interferomater for several fixed values of the absorption A. Following a suggestion of Jacquinot, the abscisse is the finesse N, which as discussed above, is proportional to the resolution or to the square root of the contrast.

For some years it has been standard practice to make use of evaporated metal films for the reflecting layers, especially of silver and of aluminum. In this case, one cannot hold A to a fixed value while R is increased by increasing the thickness of the deposited layer. R, E, A, and haves T_m are thus fixed functions of the thickness. This variation has been investigated or discussed by:

Fahry and Buisson, J. de Physique 2, 189 (1919). (For spiritured silver.)

John Strong, Procedures in Experimental Physics; silver and aluminum in

Goos, Zeits f. Phys. 100, 95 (1956).

the red and blue.

S. Tolansky, Proc. Roy. Soc., 654 (1946); Physics, 649, December (1946); Nigh Resolution Spectroscopy, Methama and Company, London, 1947. S. Dufour, Thesis, Paris, 1950.

The values the various investigators find for their practical evaporated layers are not always the same. This is partly due to the fact that the technique of evaporation, especially the purity of the metals,

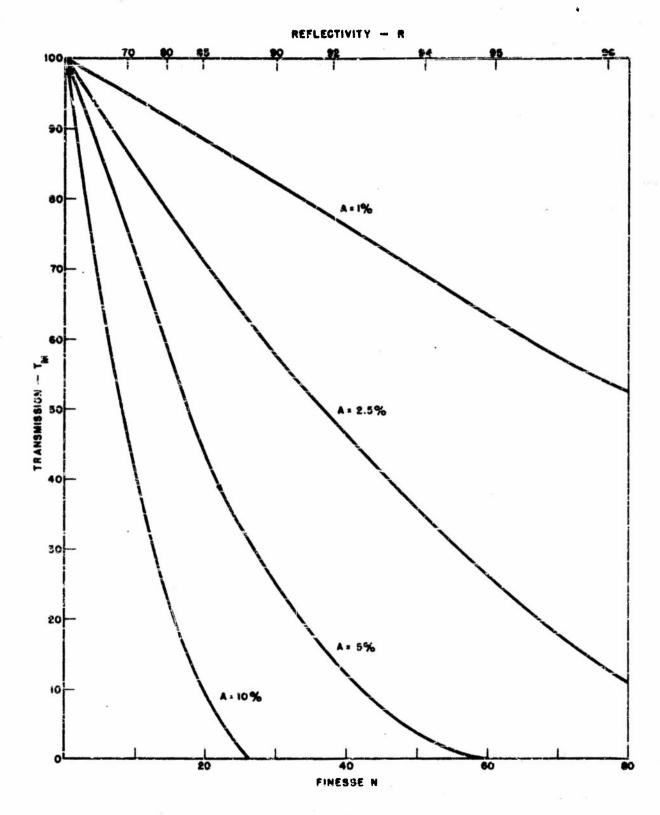


FIGURE 8

has a considerable effect. However there seems to be general agreement on the values in the region of charply falling interferometer transmission. This is plotted in Figure 9.

It is further elmost universally agreed that silver layers rapidly termish, suffering in both finesse and transmission. This can beggen in a matter of days. One case measured accurately by Dufour is above on the fiture.

At mevelengths shows 4000 angetrone silver is considerably superior to aluminum. However silver has a transparency around 5000 angetrone, so that aluminum is much better in this region. Aluminum in the red is shown in the figure.

In recent years a number of investigators have begin to make use of a new type of reflecting layer consisting of a series of layers of dislectric of alternating high and low refractive index. The history of this development is somewhat hard to trace, since a number of industrial, university and government laboratories all over the world have had extensive activity in the various applications of these sultilayer dislectrice, and few bothered to publish much. One of the most complete studies with reference to these multilayers for Fabry-Farot interferencery is contained in Dufour's thesis cited above. The general theory of multilayer interference was published by Polster¹⁰. Being Associates, Inc. has had considerable experience in this field, and some of this work is published in references cited above, and in a report on another contract.

^{10.} E. D. Polster, J. Opt. Soc. An. 39, 1038 (1949).

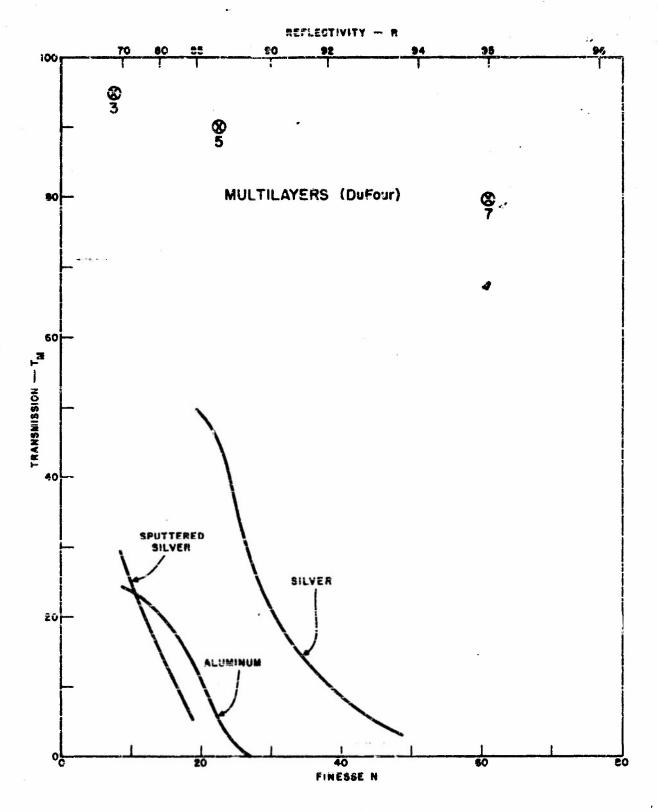


FIGURE 9

Our first experiments with the pressure tuned interferences were made with a pair of silver conted quarts plates which happened to be available. However, it was immediately recognized that the increased fineses and transmission which could be obtained with the multilayer reflectors would be vital to this difficult problem of resolving I and D light. We had ordered some especially flat quarts plates from Halle in Berlin for this job, and when they arrived, Messre. 5. Harr and 8. Sulliven of our evaporating department initiated some experiments to perfect the technique of multilayer rebry-Perot plates. The first experiments were on ordinary glass plates.

dex, there will already be a considerable reflection described by Frescal's equations. If further one stacks several of these layers together, each with an optical thickness of one quarter wave of light, the several reflecting beams will superpose in phase and the reflectivity wary rapidly experiences unity. Of course a given thickness will be a quarter wave for only a single wavelength. Actually the general theory, checked by experience shows that the reflectivity is high and approximately constant for several hundred angetross each side of the design peak. Approximate forester, valid for any master of layers over 5, were suggested privately to the writer by Jacquinot. These are:

$$1 - R = \frac{kn_0 n_g}{n_L^2} \left(\frac{n_L}{n_g}\right)^{Q+1}$$
 (15)

where no, nor no and no are the indices for air, glace, the high and low

dislociries respectively, and q is the total number of layers, and

$$E = \frac{\pi\sqrt{E}}{1-R} \approx (1.7)^{Q+1}$$

when $n_g = 1.52$, $n_L = 1.55$, $n_g = 2.5$.

Cur best results were obtained with zinc sulphide and an equal mixture by molecular proportions of lithium, aluminum and sodium fluoride. This mixture appeared experimentally to give layers of improved stability and uniformity. Pure sodium fluoride is somewhat soluble in \$\tilde{\pi}_0\$ and is susceptible to have. Eryolite, a mixture of 5 sodium to 1 aluminum, appears to be popular among some workers in the field. The best estimated indices from handbook values and from observation of the peak locations when the external is used as a spacer layer in a narrow band interference filter are:

l-l-l mix	1.36
Fayolite	1.39
Za18	2.38

These are averaged over the visible. The dispersion is not large in this region. Pure sodium and lithium fluorides were reported by Kohls¹¹ to have indices of 1.325 and 1.39 respectively.

Stone and Jeskins report that it is of considerable importance to purify 2nS by a pre-evaporation which presumably being off any free sulfur. In deposition almost all workers determine the thickness by observing the rise and fall of transmission through a monitor plate inside the evaporating chamber as the thickness is increased. Apparently sums 11. Bohls, Ann. d. Phys. 129, 455 (1957).

manifer with a fresh surface for each layer, while others observed the total layers evaporated to this point. The writer has suggested that higher accuracy in this monitor determination might be achieved by using a monitor light at a different mavelength than that for which the multilayer is designed, where one could stop deposition at the maximum rate of charge of transmission rather than on a bread peak. This has not been tried carefully as yet, however.

Applying the Jacquinot approximation to our assumed indices, we obtain the following values for the reflectivity to be expected for 5, 7 and 9.

Number of Layers	Reflectivity	Fineses
3	8 2.9	22
7	95.4	57
9	98.4	190

A five layer Fabry-Perot plate pair would be perhaps a factor of two more transparent than silver, but might be somether more difficult to make. The ? and 9 layer reflectors are considerably superior to silver of the same reflectivity. The finesse has become so high here that the flatness of the plates will be the chief limitation to resolution and contrast. Actually it is not desirable to carry the multilayers to far, as it becomes increasingly difficult to maintain uniformity in thickness and good cytical quality over so many layers. It was decided to go ahead and cost the Halle quartz with 9 layers for use in the interferementer.

It was soon apparent that it would be desirable to have accurate means for testing the reflectivity of Pabry-Perot plates in order to

determine whether the theoretical expectations of the multilayer theory were being achieved. Also, it was of great importance to be able to measure reflectivity and flatness independently of each other, in order to determine which was the principal limitation on instrumental resolution and contract.

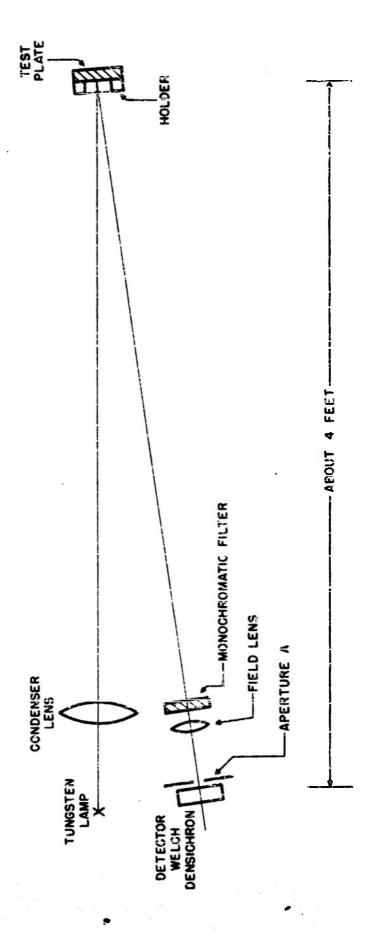
Our first brute force approach to this problem is shown in Figures 10 and 11. A tungstem lamp, narrow pass filter, Welch photometer head and mount for reflecting plates were built into a rigid system as shown. This enabled reflecting plates to be interchanged in the mount so reproducibly that the relative ratio of the reflectivities of two plates could be determined to about one percent accuracy.

Again Jacquinot in the course of his very fruitful visit to the laboratories suggested an improved apparatus. This was described by Giaccap. The instrument constructed on this project is shown in Figures 12 and 13. The procedure is as follows:

- a. Heasure deflection with no plates equals K. (Usually the Welch Densichron is just set for zero density under this condition.)
- b. Then the deflection with the first plate will measure $\mathbf{RT}_{1}(1-R_{d})$. R_{d} is just the correction for reflection at the air glass interface, about 4β .
 - c. The second plate alone deflects $KT_{e}(1 R_{e})$.
- d. Then with both plates adjusted in Fabry-Perot parallelism, the deflection is

$$ET_1T_2 = \frac{(1-R_0)^2}{(1-R)^2}$$
 where $R = \sqrt{R_1R_2}$

12. Pierre Giacumo, Scrupt. rend. December 1952, p. 1627.



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CONDENSER IMAGES SOURCE ON FIELD LENS. FIELD LENS IMAGES APERTURE A ON TEST MIRROR. APPARATUS FOR DIRECT REFLECTIVITY MEASUREMENT

FIGURE 10

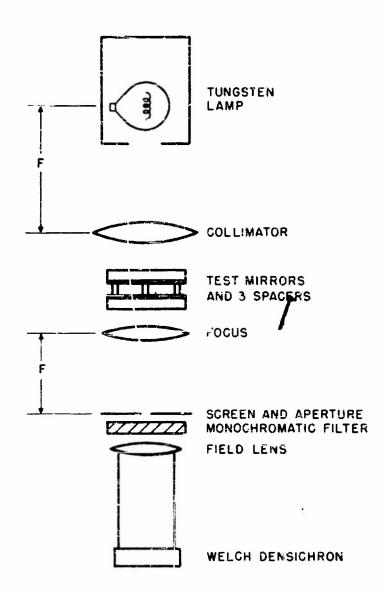


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are to war expert. For March 1



APPARATUS FOR MEASURING I-R

FIGURE 12





La Dan Branco and Carlotte for the same

Thus solving,

$$\bar{r}_g = \frac{c}{a(1 - \hat{R}_d)}$$

$$1 - R^2 = \frac{bc}{ad}$$

CT

$$R_1 R_2 = 1 - \frac{bc}{cA}$$

or if an approximate value for H is known,

$$1 - R = \frac{bc}{ad} \left(\frac{1}{1 + R} \right)$$

The advantage of this method over the former is that it directly measures low rather than R. Although the other equipment could measure R to 15 accuracy, the accuracy in 1-R could become quite large, of R was, for example, about 95.

The result for the value d when measuring through both plates is obtained resulty from equation (2) for the transmission of the Pabry-Perot. The condition is that the thickness of separation and the width of the wavelength band in the white source must be such that the source emission does not appreciably change over the free spectral range of the interferometer. Then one is justified in integrating the intensities of equation (9) over one whole order, obtaining the result given. The condition on the thickness e and the source width AA is

4

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For A = 50 angetroms at Mx, this prescribes 6) > 256.

Adequate spacers were found by outting small chips out of some callulaid material. The instrument was arranged vertically so that one plate could be laid down, spacers on it, and the second simply laid over the spacers. It is not at all difficult to determine whether the two plates are sufficiently parallel. A small specture just larger than the filement image is laid on the stage and moved until the image completely passes through it. If there is a wedge in the plates, sailtiple images will be observed. When the plates are parallel, all these images will pass through the aperture, which is all that is required.

Although the method only gives the product of R_1R_2 , with three plates all three R's can be determined absolutely by measuring the products for all three pairs.

By meens of this apparatus, the reflectivities of our 5, 7 and 9 layer plates were measured to be 87, 95 and 97 percent respectively, values in good agreement with the theoretical prediction.

Scar of the experimental multilayer plates are shown in the photograph near the instrument for determining reflectivity.

The resolution and contrast of the Malle quarts plates, as limited strictly by the reflectivity, were considerably greater than was actually obtained in the Fabry-Perot fringes in Ex light. It was time evident that the effect of plate flatness and source line width would have to be investigated.

5. EXTRECT OF YLICENSES, SOURCE VIDER AND APERTURE SIZE

so far the consideration of the relative shape, resolution and contrast of the fringes has been limited to the effect of the reflectivity of the plates. It has been assumed that the plate thickness t and the wavelength A were guite constant. When these ideal conditions are not fulfilled, the distribution of plate flatness defects and of wavelengths in the source may contribute to fringe broadening and lose of contrast. The effect of the size of the circular sporture before the photocell in also to be considered. It is of interest to obtain the expression for the absolute light energy transmitted through the entire instrument, both at the peak and at the minimum of the fringe, for noise considerations.

In any optical system the total light transmitted is given by the product of source brightness by the area and the engular sperture of the system, multiplied by any transmission factor. The exact expression for the absolute energy transmitted by the Pahry-Perot instrument considered here is thus given by

$$E = \iiint_{C} J_{C} dC \cdot A(t) dt \cdot 2x + C \cdot I(R, C, t, \bullet)$$
 (16)

This triple integral as derived in the following way. A_0^{-} is the spectral brightness distribution of the source. A_0^{-} is thus the total brightness in the range A_0^{-} . It is convenient to use wave number rather than wavelength as the variable because the expression for order number n is linear in wave number, but inverse with wavelength. A(t) dr is the area of the Febry-Perot plates having thickness between t and tedt. The distribution function A(t) thus describes the departure of the plates from

exact flatness and parallelism. Excés is the angular specture of the light on it peaces through the interferometer. That is, it is the solid angle subtended at the plates by the angular ring in the image plane of the fringes between 0 and 0.000. In practice an aparture is placed in this plane and the light passed through this specture to a photocell. Consideration is limited to centered annular rings in order to have to deal with only a one dimensional integral for the solid angular specture. In the experiments the aparture was always a simple circular opening, but the results are equally simple and general for any annular ring extending from 0, to 0, the one dimensional integral in 0 being taken between these limits. The final factor I is the transmission of the interferometer, which has been previously given in equation (8), but which is now rewritten in terms of the variables n and 6.

$$R(R, \delta, t, \phi) = R_{\rm H} \frac{1}{1 + a \sin^2(\pi a)}$$

where T is nearly one for multilayer reflectors (no absorption),

$$n = \frac{4\pi}{(1-R)^2}$$

$$n = 8\mu t \delta \cos \theta \approx 8\mu t \delta (1 - \frac{1}{2}\theta^2)$$
 %<1

There are four more or less sharply peaked distributions in this triple integral, the wavelength distribution, the flatness distribution, the reflectivity limited frings shape and the sperture transmission, which can be considered as an ideal square wave distribution in 0. The width of each of these distributions makes its contribution to the overall fringe

width. It is the purpose of this section to consider with some right these relations.

by the Doppler broadening. A few remarks on the flatness distribution seem to be in order. There are certain simple types of thickness defect which lend to simple analytical distributions which can be easily treated. First, there may be random roughness of the plates on a rather small minroscopic scale, due to the roughness of the polishing process. It seems reasonable to assume that these defects have a Gaussian distribution. A second simple type is a wedge, i.e., a departure from parallelism. If the aparture which is utilized is square, it is obvious that the distribution of thickness in the wedge is a square wave. These two types of distribution are indicated graphically in Figure 14.

It is perhaps surprising to find that the distribution for a contered spherical error, which is a common flatness error, when used with a circular sporture, is likewise a square wave, thus identical in shape to the previous case.

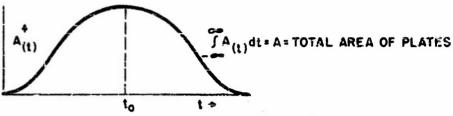
It is of considerable convenience to make a change of the variables of integration in the general formula (16), so that thickness, were muster and angle 3 will be reduced to the same dissensions and can be considered on a common level. The order number n of the interference is an ideal varuable for this purpose. We write

$$t = t_0 + t'$$
 $6 = 6_0 + 6'$
(27)

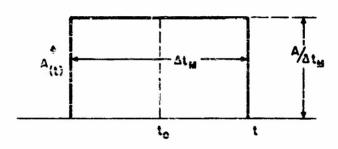
FIGURE 14 DISTRIBUTIONS OF SURFACE DEFECTS

I. GAUSSIAN DISTRIBUTION OF MICROSCOPIC ROUGHNESS

A(t) dt = FRACTIONAL AREA HAVING THICKNESS BETWEEN t AND t -+ dt



2. WEDGE OR SPHERICAL DEFECT



a WEDGE



b. LENS DEFECT



AREA OF PLATE HAVING THICKNESS BETWEEN y AND y + dy = $2\pi x dx = 2\pi R dy$ FOR SINCE y = $\frac{x^2}{2R}$, Rdy= xdx

R = RADIUS OF CURVATURE

NORMALIZATION: $A = \pi x_m^2 = 2\pi \pi \Delta t_m$ $A(t) = \frac{A}{\Delta t_m}$

where $t_{e^{\pm}} \leqslant_0$ and θ_e are the average values of their respective distributions, and the primed variables are the small deviations. Then the expression for n, to the first order in the small deviations, becomes

$$n = 2\pi (t^0 + t^1)(t^0 + t^1)\left\{1 - \frac{3}{2}(t^0 + t^1)^{\frac{1}{2}}\right\}$$

$$n = 2\pi (t^0 + t^1)(t^0 + t^1)\left\{1 - \frac{3}{2}(t^0 + t^1)^{\frac{1}{2}}\right\}$$

which we write as

introducing the variables n as the equivalent order numbers of the various deviations from n. That is

$$n_{o} = 2\mu t_{o} t_{o} (1 - \frac{1}{2} t_{o}^{2})$$

$$n_{e} = 2\mu t_{o} t_{o}^{2}$$

$$n_{e} = 2\mu t_{o}^{2} t_{o}^{2}$$

$$n_{e} = 2\mu t_{o}^{2} t_{o}^{2}$$

$$n_{e} = 2\mu t_{o}^{2} t_{o}^{2} t_{o}^{2}$$

$$(18)$$

This clarge of variables is then introduced in equation (16), giving

$$R = \frac{2\pi}{n_0} \int_{n_0} \int_{n_0} \int_{n_0} J(n_0) dn_0 \cdot A(n_0) dn_0 \cdot dn_0 \cdot I(R, n)$$
 (19)

It is essued that the distributions J and A have been reexpressed with n as variable, maintaining $J(f)df = J(n_f)dn_f$ and

A(t)ds = $A(x_t)$ ds. This requires

$$A(a_t) = \frac{1}{a_t G_a} A(t)$$

The triple integral still seems formidable. The integration over sporture can be reserved until last, since the distribution is always perfectly square wave. Thus for the distribution of energy in the plane of the image, one has

$$\mathbf{x}_{\mathbf{n}} = \frac{\mathbf{E}}{d\mathbf{n}_{\mathbf{n}}} = \frac{\mathbf{g}_{\mathbf{x}}}{\mathbf{n}_{\mathbf{0}}} \iint J_{\mathbf{n}} d\mathbf{n}_{\mathbf{0}} \cdot \mathbf{A}_{\mathbf{n}} d\mathbf{n}_{\mathbf{0}} \cdot \mathbf{I}(\mathbf{n}, \mathbf{n})$$

One interesting conclusion may already be stated about the distribution in the sperture. Since I(n) is periodic and repeats exactly with such integer, the total energy pasced through my annular ring is dependent only on the fractional value of n at the two edges of the aperture. That is, in the multi-ringed pattern of Fabry-Perot fringes, the total energy in each ring is the same, although the dismeter and width of the successive rings very as one moves out from the center. This result is simply due to the fact that the area in the image plane between half maximum points, for example, is the same for all rings.

If the width of the 3 and A distributions is small compared to the width of I(n), we obtain some rightnessly the results stated in an earlier section for the reflectivity limited case. The value of I(n) is that sensibly constant over the range of variation of n_{ℓ} and n_{ℓ} , and the integrals are separable.

where

$$A = \int A(n_t) dn_t = \int A(t) dt = \text{eres of interferometer plates}$$

$$B = \int J_6 d6 = \text{integrated line brightness}$$

Now it does not matter how narrow these distributions may be, so long as they are narrow with respect to I(n), the fringe energy will still have the same shape and absolute value. It is further clear that the specture may be opened up, with increase in the total peak energy and without decrease of contrast, until the peak energy begins to fall off. That is, the proper specture is the one which just about subtends the half width of the final fringes. The rigorous integral of Ion for this specture gives $\frac{\pi}{4} \delta n_R$ and the contrast is reduced from its peak value by $\frac{\pi}{4}$.

$$n_a = +1/2 \, \delta n_R$$

$$\frac{1}{1 + m \, \sin^2(\pi n_a)} \, dn_a$$

where
$$n = \frac{4R}{(1-R)^2}$$

$$\delta \alpha_{\widetilde{R}} = \frac{1-R}{\pi \sqrt{R}}$$

This integral is found in Pierce's Tables, formula 314. The result is

$$\frac{1}{\pi\sqrt{n+1}} \left| ten^{-1} \left(\sqrt{n+1} ten \pi n_{\alpha} \right) \right| + \frac{1}{2} ten_{R}$$

$$\frac{1}{\pi\sqrt{n+1}} \left| ten^{-1} \left(\sqrt{n+1} ten \pi n_{\alpha} \right) \right| - \frac{1}{2} ten_{R}$$

For high values of R, $\frac{\pi}{2} \delta n_R = \frac{1-R}{2\sqrt{R}} \ll 1$, and ten $\theta \gg \theta$. Hence the integral is given fairly closely by

$$\delta n_R \frac{2\sqrt{5}}{1+R} \tan^{-1} \frac{1+R}{2\sqrt{R}}$$

From R = 80% to R = 100%, $\frac{1+R}{2\sqrt{R}}$ is very close to 1 (within .01). The angle whose tangent is one is $\pi/4$ radians. Hence

$$+\frac{1}{2}\delta n_{R}$$

$$-\frac{1}{2}\delta n_{R}$$

$$-\frac{1}{2}\delta n_{R}$$

Thus the absolute available energy and contrast for monochromatic light and flat plates in

$$\mathbf{E} = \left(\frac{2\pi AB}{n_0}\right) \frac{\pi}{h} \delta n_R = \frac{\pi^2 AB}{2n_0 E_R}$$

$$\delta n_a = \delta n_R$$

$$(\delta n_R = \frac{1}{B_R} = \frac{1 - R}{\pi \sqrt{R}} \approx \frac{1 - R}{3})$$
(20)

Thus one can improve the contrast, but at the expense of energy by increaseing the reflectivity toward unity. Eventually of course one sawt reach the point where the reflectivity width $\delta n_{\rm R}$ is less than either

the source width or the flatness defect. It is of interest to investigate this case. By a similar argument to the one used above, it is evident that only the wider of the two distributions will be of interest. In our experiments, for example, it is fairly conclusive that the source width is the limiting factor, about 1/12 order, the plates being flat to at least 1/60 order.

The required integral for the total energy is now

$$E = \int_{\mathbb{R}_n} \mathbb{R}_n dn_n = \frac{2\pi}{n_0} \int_{\mathbb{R}_n} \int_{\mathbb{R}_n} J(n_0) dn_0 I(R, n) dn_0$$
aperture aperture $\left[\frac{1}{n_0} \right]$

We consider the case when the width δn_R of I(R, n) is much less than the width δn_ℓ of J(n_\ell). We integrate the expression in brackets for two choices of the sperture, first at the peak, then at the minimum midway between fringes. First, at the peak, I(n) varies rapidly while J remains constant. Thus the bracket integral becomes

But since I(n) has a peak of unity end a bull width $\operatorname{On}_{R^{\prime}}$ this last integral has the approximate value $\operatorname{On}_{R^{\prime}}$. The fringe shape is thus J, the source shape. This time the aparture might as well be opened to the half width $\operatorname{On}_{R^{\prime}} = \operatorname{On}_{\ell}$ of the source, and the peak available energy has the approximate value

$$\mathbf{E}_{\text{peak}} = \frac{2\pi A}{n_0} J(n_0) \delta n_0 \delta n_R = \frac{2\pi AR}{n_0} \delta n_R$$

This is the same as obtained in the previous cape.

At the minimum of the fringe I(n) changes very slowly, so the bracket integrates to

and $I_{min} = 1/6_R$, $6_R = 16aal$ reflectivity limited contrast and the total energy for exerting $2m_L = 8m_6$ is

Bases, the effective contrast is

$$e = \frac{R_{\text{peak}}}{R_{\text{con}}} = e_{R} \frac{\delta u_{R}}{\delta n_{C}} = 0.4 R_{R} u_{C}$$
 (21)

Thus the reflectivity contrast is not obtained when $\delta u_{\mathbf{R}} \leqslant \delta \mathbf{n}_{\mathbf{f}}.$

Conclusions:

- (1) The total energy in each Fabry-Perot ring is the same.
- (2) The total peak energy is in absolute value approximately

$$\frac{543}{n_0} \, \delta n_R = \frac{5}{3} \, \frac{AB}{n_0} \, (1 - R) \tag{52}$$

B = source brightness, A = plate eres, $\frac{56n_h}{n_h}$ = effective solid engle.

(5) If the fringe width is purely reflectivity limited, the contrast is

$$E = 0.4 \, E_R^2$$
; $E_R = \frac{e \sqrt{g}}{1 - k}$ (25)

(4) If the fringe is source width limited

$$a = 0.4 \text{ MyH}_{c}$$
; $a = \frac{\Delta c}{8c} = \frac{1}{8c \cdot 5c}$ (84)

If flatness limited, replace He by He.

(5) Given a source line width, the optimum resolution, energy and contrast are achieved when the plates are at least the same flatness or better, and the reflectivity and sporture setting satisfy

$$\delta n_p = \delta n_a = \delta n$$

6. TESTANG THE FLATNESS OF FARRY-PEROF PLATES

At the start of this project there were a number of quartz plates available which had been prepared by Mr. Wagn Hargbol of Arlington, Massachusetts. Mr. Hargbol reported that at that time he had no facilities for testing flats to better them 1/50 frings of light, but that these plates were at least this flat. Inquiries were made of numerous sources of quartz flats. Hilger in England and Halle in Berlin thought that their plates were flat to at least 1/75 frings. Two plates were ordered from Halls. The Hilger plates were considerably more expensive.

It was considered desirable to be able to test flats in our laboratories, in order to test the claims of various producers, in order to determine to what extent the Fabry-Perot resolution would be limited by flatness, and further because it was felt that the local simplier, Mr. Marghol, could make plates as well as anybody if he only had some means to test the progress of his work.

There are a number of methods for testing flats which have been described in the literature. Almost all are based on some form of multiple been interference fringes produced when the plates are coated with highly reflecting surfaces and mounted with a very slight wedge between them. Two simple optical systems for observing these fringes are sketched

in Figure 15, one in transmission and one in reflection. In transmission narrow light fringes on a dark background are produced, and in reflection narrow dark fringes on a light background. The transmission method was considered simpler and made use of.

The first tests were made with silver costed plates. The fringer were too wide for accurate determination of flatness. Typical results are indicated in the photograph of Figure 15⁸. The width of these fringes second to be about 1/10 the fringe separation. These is little observable departure from straightness, but it would be hard to detect better than 1/20 fringe with such wide fringes.

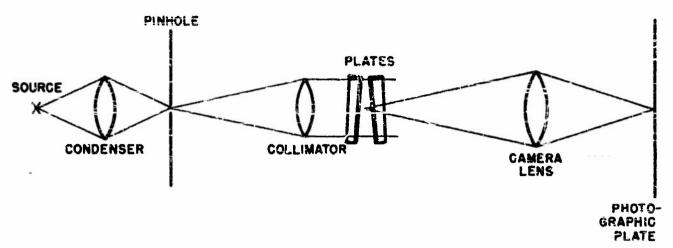
One of the best discussions of the conditions to be not to eableve narrow fringes for plate testing purposes is found in Tolansky's "Multiple Beam Interferometry". It requires multiple beam interference to produce fringes narrow with respect to their separation. The fringes are produced when

DA - But cos 4

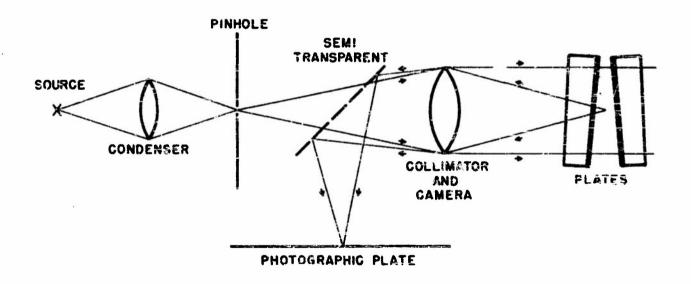
There are three variables (sounding μ constant), t the plate separation, θ the angle of observation and λ the unvelength of the illumination.

If t and h are hald constant we have fringes of equal inclination. These are the Fabry-Perot fringes, with which most of this report is concerned. The transmission varies with 8 and the fringes are therefore formed at infinity.

A Some of the early work on flatness testing at these laboratories is reported in detail in an earlier technical report, reference 6.



A. PLATE TESTING IN TRANSMISSION



B. PLATE TESTING IN REFLECTION

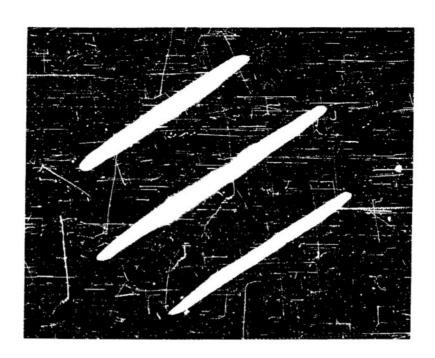


FIGURE 16.



FIGURE 17.

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If \$ and \$ are held constant the same fringe-intensity distribution results (equation 6), but the variation with t is observed. These are called fringes of equal thickness and the fringes are focussed approximately in the plane of the plates. These are the fringes then that are useful in observing the flatness of plates.

The third type, with t and 0 constant, λ variable, Tolersky calls fringue of equal chromatic order.

Actually, in the pressure tuning method, the fourth variable μ is used.

There are four conditions which must be satisfied in order to obtain the sharpest fringes.

(1) The fringes cannot be narrower than the reflectivity limit discussed above

$$\delta n_R \approx \frac{1-R}{3}$$

(2) The absolute separation between the plates must be kept very small. This is because of the wedge introduced to produce the series of fringes across the face of the plate. As a result the phase change for each successive bess in the multiple interference is not the same, as in the perfectly parallel case. By detailed consideration of this effect follows shows that the absolute plate separation must be less than

$$t < \frac{1.5(\delta n_{K})^{3}}{\lambda x^{2}}$$

where X is the number of fringes per centimeter across the wedge. When there is no wedge (Fabry-Perot) then t can be large (i.e., X is small).

(5) Above it was stated that 4 is to be constant, i.e., parallel light. Here detailed consideration shows that the collination must be at least

(4) The source width must of course be less than the reflectivity width.

In the first attempts the principal difficulty was with the poor reflectivity of the silver. This was remained by costing the Hallo plates with nine multilayer films, having a reflectivity of about 97%. This gives a theoretical fringe width of .01 fringes. The other three conditions were set, and the fringes shown in Figure 17 were obtained. These are by the optical arrangement of A, Figure 15.

These fringes were quite sharp. It was estimated that the width was at least 1/80 fringe. On the original plates fine microscopic unggles can be detected which are presumed to be the microscopic roughness of the plate. The central portion of the plate seems to be flat to around 1/80 frings. There is a pronounced rounding of the fringes toward the adge of the reflecting layers, which were about 1.4 inches dismeter on the 2 inch plates. It is not known whether this rounding off is due to lack of plate flateess or whether it may be caused by non-uniformity of the expertated sultilayers. Further study of the uniformity of laying down nine dielectric layers would be required. It was evident however

that for the central portion of the plates, the flatmess would be less of a limitation on the hydrogen-deuterium enalysis than the actual width of the hydrogen lines in the source.

7. DEPROVING WINGTRONIC SIGNAL-TO-MOISE

The work on the testing of flatness and reflectivity in the PolicyPerot plates was now convincing that the resolution was limited only by
the actual source width. What was being observed was the Doppler broadened
fine structure of the E-C line, the source broadening being so large that
the seven actual fine structure components were observed marely as two
broad lines just berely resolved, that is, highly overlapped.

At this time recorder traces as indicated in Figure 18 were being obtained. The contrast, that is, the ratio of the peak H intensity to the minimum H-fringe intensity at the D position, was about 200. (Dark current and room stray light were of course subtracted out.) The cample for Figure 18 was prepared to approximately a deuterium concentration of 0.5%, and the ratio of the D-light to the H-light was approximately this value.

It was apparent that further significant gains in pushing the concentration sensitivity to 1 wer deuterium concentrations would be obtained only by improving the electronic technique of reading the D peak in the presence of the E-berkground. How in such situations, it is obvious that the background itself, if it were constant, is no obstacle to infinite sensitivity for it can be merely subtracted out. What will ultimately limit the sensitivity is the fluctuations in the background. Theoretically, if one could eliminate all other sources of noise, the

R The fine structure and Doppler broadening of Ex and E are discussed in the preceding Technical Report under this contract, dated July 1, 1953.

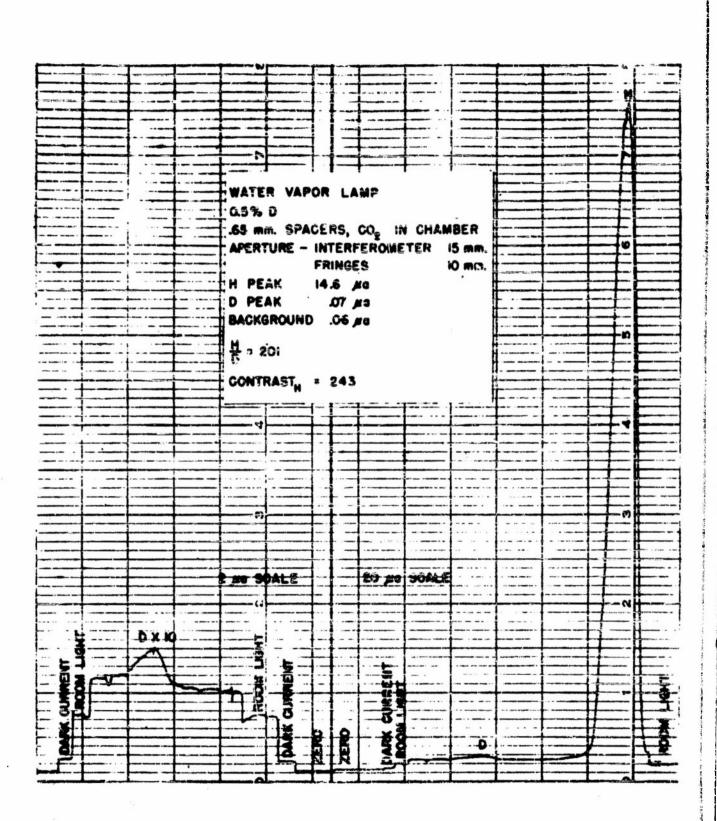


FIG. 18

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ultimate limit would be set by the shot effect fluctuations in the electrons emitted from the cathods of the photomultiplier. The magnitude of this effect is well known.

The rae fluctuation noise in the cathode current is given by

$$(\nabla I^F)^{1.38} = \left(\frac{1}{r^3}\right)_{1/5}$$

where I_R is the average cathode current, e the charge of the electron; and T the time constant of the overall detector—indicator. The current I_R can of course be expressed as the product of the available light flux F by the absolute sensitivity S of the photo-cathode, and besides these two factors, only the time constant influences the noise level. Thus one can increase the signal-to-noise ratio only by increasing the light flux, the photo-cathode sensitivity or the time constant.

$$\frac{S}{S} = \frac{I_{k}}{\Delta I_{k}} = \left(\frac{I_{k} \gamma}{e}\right)^{1/2} = \left(\frac{F_{k} \gamma}{e}\right)^{1/2}$$

In the photomultiplier the gain of the dynode stages is simple emplification of the noise and signal by the same factor \tilde{u} , so that the signal-to-noise is not changed in the succe current. However since the cathode current is not readily measured, it is perhaps convenient to record the noise equations relating to the anode current I_A .

$$(\Delta I_A)_{\text{TEMS}} = G(\Delta I_K)_{\text{TEMS}} = \left(\frac{G I_A e}{I}\right)^{1/2}$$

$$\frac{8}{h} = \frac{I_A}{\Delta I_A} = \left(\frac{I_A I'}{eC}\right)^{1/2}$$

[#] See, for example, R. W. Engstrom, J. Opt. Soc. Am. 37, 420 (1947) or R. O'B. Carpenter, J. Opt. Soc. Am. 40, 225 (1950).

This has the same numerical value as the expression obtained above, but it is usually laborious to measure the photomultiplier cathode nurrent and absolute consistivity. With the last expression one must estimate the gain G. A rough value can be obtained from the BEA handbook, but too such reliance should not be placed on it because of the tremendous variation in gain from sample to sample:

The above expressions for S/N assume that the entire current somsists of the signal. When there are extraneous currents added to the signal, such as photomultiplier dark current or room stray light, the above expressions for S/N are to be multiplied by the ratio of the signal current to the total current.

Therefore the photomultiplier and recorder system were critically investigated in order to determine the limiting noise factors and the ultimate concentration sensitivity which could be obtained in hydrogenedeuterium determinations. The hope was to be able to reach the detection of the deuterium existing in ordinary water, namely .Olf percent.

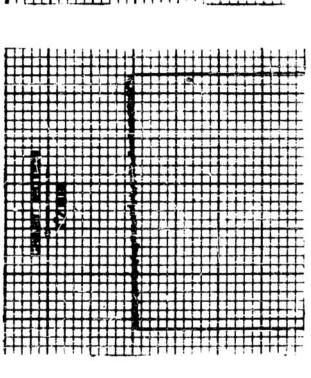
A first difficulty appeared because the maximum scale sensitivity of the Leeds and Northrup Speedomax being used was 2 micrompares full scale. This strip chart recorder is "dend" for deflections less than about 0.2 percent of full scale. With this scale and 1000 volts on the photomultiplier it was just berely possible to observe fluctuations in the trace. Therefore it was desirable to increase the sensitivity of the recording system. This could be done in two ways. The photomultiplier voltage, and hence its gain G, was increased. Soon the safe limit of 1850 volts was reached. The imput circuit of the Speedomax was investigated to see whether

its sensitivity could be increased. It was found that a simple change would give us a factor of 2, or a 1 microsspere full scale setting. It impeared that any further gains would be tiles—community in that the entire coulding and damping input network of the recorder would have to be completely redesigned.

Although we would have preferred to be able to go further with the increase of gain, considerations of time and expense precluded this, in view of the fact that with these simple changes fairly considerable noise fluctuations were now detected. These are shown in Figure 19.

Experse on the 1 microsupers scale. This current was derived from a dry bettery and resistence and shows that there is no noise introduced by the recorder. The recorder also showed no dead anot or overshoot within 0.25 full scale. The next trace is a light signal from a taugeten lamp operated from a storage bettery, which is known to give a highly stable and noise-free light flux. The gradual drift in current may be fatigue of the photocell. At any rate, the noise fluctuations are quite large. The succeeding trace shows the reduction of this noise when the time constant of the recorder is increased by simply connecting two 8 µf oil expections in parallel across the input leads. The effective input resistance of the recorder is about 70,000 chas so that the time constant can be estimated from the R6 value. With no external expectance the time constant is about 1 second, being limited both by cone internal capacity and by the inertia of the motor-pervo system.

FIG. 19A
0.75 ## FROM DRY BATTERY.
DFAD SPCT TESTS



0.75 ag from photomultiplier. 1250 v on photomultiplier and Battery operated tungsten lamp

FIG. 198

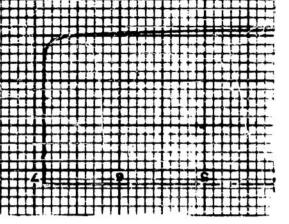


FIG. 19C 16 Jf Adden Across IN-Put for 198

TESTS ON I Ma SCALE OF RECORDER

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8. FIRAL EVALUATION OF PERCONANCE

Three other sources of noise were noted at various times. Variations in photosultiplier gain due to line voltage fluctuations were prectically eliminated by use of a Sorenson voltage stabilizer.

Stray room light became significant when attempting to measure low concentrations. Most of the light reaching the photomultiplier seemed to be from reflection off the vacuum chamber hobing the interferometer. The uncented light reaching the photomultiplier was reduced to a negligible amount by putting the photomultiplier in a long box with a variable iris just in front of the interferometer focusing lens. This iris also actuat the limiting aperture for the Fabry-Perot interferometer.

The microws a exciter used with the lamp was also operated from a voltage regulator but the magnetron output still had an occasional dip which results in a sharp intermittant pip on the recorder trace. This does not interfere seriously with the measurements but is certainly not desirable.

After the noise reductica work a sample of approximately 0.05% D was prepared.

Two limitations of the quickly assembled equipment used for these preliminary studies becase apparent. First the interferometer spacers are 0.65 mm thick. This gives a difference in order between Mx and Dx of 0.54 order instead of the difference of 0.50 order assessary to put D maxisum at the H minimum. For an 0.055 concentration of D and a contrast of 200 the ratio of D max to H_{min} is

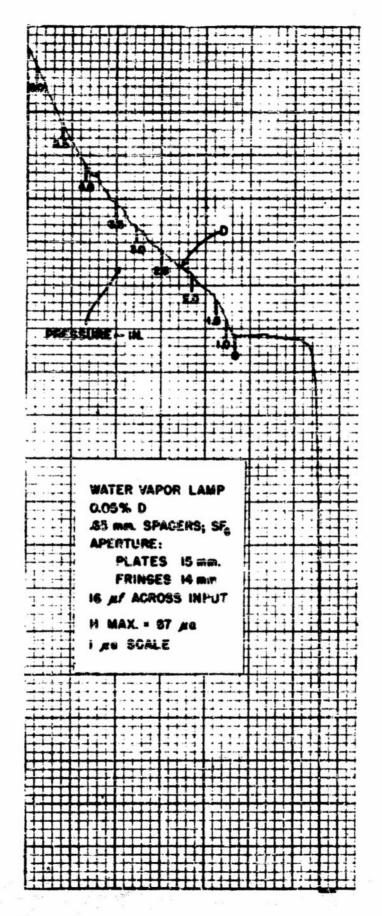
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Since the H fringe increases by a few percent over its adminum at .54 the D maximum is seen as just a bulge on the side of a sloping line. Since the control of pressure change is rather crude the recorded traces are not linear with pressure (or order). The rate of change of pressure can cause an irregularity in alope which may be discouragingly similar to the D maximum as in Figure 20.

closely related to this difficulty. The parallelism of the plates is the only factor affecting the contrast which is likely to change in the short time of a few days or weeks. Besides the obvious effect of making the D_{max}/H_{min} ratio smaller it also changes the shape of the fringes apparently in a way that is not helpful. The trace in Figure 20 was for a contrast of about 130 and a D concentration of 0.0%. The trace in Figure 21 is for the same conditions except that the interferometer had been adjusted so that the contrast was 210. Here the D_{max} can be seen quite easily and D_{max}/H_{max} is about 0.0%. However the elope of the E background makes the measurement of the D peak rather difficult. It is definitely of great importance to obtain ball spacers of the exact required thickness. A samufacturer will probably have to be asked to make a special run.

There was not enough time to make a series of runs with natural water but one test seems to indicate that the O.OLOS D concentration is detectable with the present equipment although it commut be measured with much accuracy.



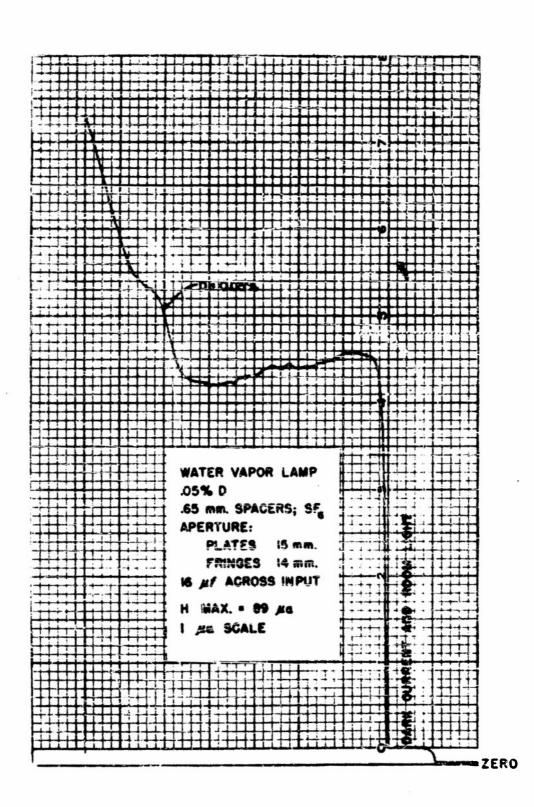


FIG. 21

9. COMPARISON OF THE LUMINOSITIES OF FAMILY-PEROR, CENTURY, PRISM AND INCLARIZATION CHARGEFUL CONTRACTOR

In evaluating a variety of dispersing instruments for use in spectrochemical analysis, one of the points of interest will be the relative light gathering power of the various instruments, sometimes called the "luminosity". This is particularly true when one wishes to measure quite low concentrations of an element in a source which is inherently not particularly bright, in other words, in an energy limited situation. The luminosity may then determine whether the electronic noise is limited by the photon shot noise, or by noises originating in the determine-emplified combination.

The general formula for the energy gathering power of any opti-

E - BANT

where B is the source brightness

- A is the area of the optical been
- m is the solid angular divergence of the been
- T is the relative transmission of the system as limited by absorption or reflection.

This formula will be applied in turn to the Fabry-Perst interferometer, a grating spectrograph, a prime spectrograph, and the polarization interferometer. Since B refers to the source, and may be assumed to be the same for competing dispersing elements, the figure of merit of the dispersing elements will be the luminosity, defined as

L - And

This quantity will have the dimensions of error.

The luminosity of any instrument will depend on the resolution which has to be achieved. It will be assumed that the problem requires a resolution $R = \lambda/\delta\lambda$, and that all the source energy is in a bent smaller than $\delta\lambda$.

1. The Fabry-Perct Interferometer

The work of calculating the sporture and lustmostly of the instrument has already been performed in the preceding section. The result for the total energy passed was

Movever in an earlier section of this report, now was found to be the resolution R of the system. Hence the luminosity of the Fabry-Ferot is given by

Note: the factor $T_M = 1 - \frac{A}{(1-R)}$ has been left out on the assumption that non-absorbing multilayer reflectors are used. For eiliver or other absorbing reflectors this factor should be included.

2. The Grating Spectrograph

For the grating spectrograph the solid angle of the grating as seen by the exit slit is

A 100 r

where $A_{\bf g}$ is the area of the grating rulings, r the engle of diffraction and f the focal length.

The transmission factor of the grating varies considerably from grating to grating, and with wavelength, order, blaze, etc., but for a properly chosen well-blased grating might easily be about 50% at H alpha.

The area of the slit is of course simply its width times its length. The width is limited by the resolution desired and the dispersion. The dispersion is given by

$$n_0 \lambda = a(\sin i + \sin r)$$

$$\frac{\partial x}{\partial x} = \frac{1}{f} \frac{\partial x}{\partial r} = \frac{a}{f R_0} \cos r$$

where r is the angle of diffraction, n₀ is the order number, a is the greating line space constant. To achieve resolution $R = \lambda/2\lambda$, we must use a slit width

$$\Delta x = \Delta \lambda \frac{dx}{d\lambda} = \frac{\lambda}{R} \frac{dx}{d\lambda} = \frac{\lambda f \ n_0}{R \ a \cos x}$$

nows can be eliminated by means of the grating formula above

$$\Delta x = \frac{f}{R} \frac{(\sin 1 + \sin 7)}{\cos 7}$$

The length 1 of the slit is commonly of the order of 1 inch.

It is limited usually by the fact that above or below the plane of the Envland circle, the estigmatic image lines became not perpendicular to the Rowland circle nor parellel to the slit. For this reason Fastic devised a type of grating mount originally proposed by Ebert with circular 15. William G. Fastic, J. Opt. Soc. Am. 42, 641, 647 (1952).

rather than straight clits, with entrance alit and exit alit symmetrical with respect to the grating, in which the alit length could be increased to 4 inches and more without ostigmeticm troubles. Hence the final result for the luminosity is

$$L = AnT = \frac{T_R A_R}{R} \frac{1}{2} (\sin 1 + \sin r)$$

This result is less than the luminosity of the Fabry-Perot by approximately 1/5f, Ordinarily one would expect about $f = 40^{\circ}$ (1 meter), $1 = 1^{\circ}$, or the grating is about 1/200 as fast. As pointed out above, astigmetims usually restricts the slit length.

3. The Prish Spectrograph

In the case of the prism the derivation will be exitted. But in the case of minimum deviation, which is a condition for maximum luminosity, the result can be written:

$$L = \frac{2A}{R} \frac{1}{I} \frac{\lambda}{n} \frac{dn}{d\lambda}$$

The luminosity is again limited by the slit length factor, and also by the final factor depending on the dispersion of index. The area $\hat{\rho}_{pr}$ in the above result is the area of one of the faces through which the light passes. Typical values for the dispersion factor $\frac{\lambda}{n}\frac{dn}{d\lambda}$ are between 0.1 and 0.01. So absorption losses are included in the above formula.

4. Polarisation Interferometer

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In the last technical report submitted under this project (dated July 1, 1953) an interesting instrument for separating H and D light was discussed; based on the dispersion of birefringence. This was

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called the polarization interferometer. The light was not actually dispersed in ordinary space, but the W and D light was polarized at right angles to each other. In this section the luminosity of this system is compared with those computed above.

This comparison is included in this report for completeness, and also because of the interesting properties of the polarisation interferometer. The interested reader may have to rafer to the detailed discussion of the last technical report for the definition of all the terminology.

If polaroids are used for the three polarizers their transmission will be about $.50(.70)^3 = 17\%$. 50% is lost in the initial polarization of the light. In order to get a high percentage polarization, fairly dense polaroids must be used, so that the transmission for the parallel exponent will probably gut exceed 70%. Hence

Other sources of loss in transmission will be reflection losses, glass and crystal absorption. These will be small if the system is immersed in oil, which will probably be necessary for temperature control purposes anyway.

The angular field of these x-cut orystal plates has been desoribed by Evans. For the eighle plates without the field widening technique the phase retardation as a function of the angle of transmission through the crystal is given by

14. John W. Evans, J. Opt. Soc. Am. 39, 229 (1949).

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$$\delta = \delta \left\{ 1 + \frac{\theta^2}{8n_0} \left(\frac{\cos^2 \theta}{n_0} - \frac{\sin^2 \theta}{n_0} \right) \right\}$$

$$= \delta_0 \left\{ 1 + \frac{\theta^2}{8n_0} \right\} \qquad (\text{at } a = 0, 90, 180, 270)$$

where θ , θ are the polar coordinates of the angle of incidence, n_0 , n_0 the ordinary and extraordinary refractive indices of the uniaxial crystal. For the Lyot-Evans type field widesed system (two plates of half thickness separated by a helf wave plate)

$$\mathcal{E} = \mathcal{E} \left\{ 1 + \frac{\theta^2}{4 \bar{a}_0} \left(\frac{1}{\bar{a}_0} - \frac{1}{\bar{a}_0} \right) \right\}$$

Hence the minimum improvement by the field widening (i.e., ratio of angles of incidence having sees (AV) is

$$\left(\frac{2n_0}{n_0-n_0}\right)^{1/2} \approx 8 \text{ for All }$$
All for calcite

18.2 for quarts

The above formulas give the deviation in retardation for the ray deviating from the normal by θ degrees. If we consider a circular cone of rays having a maximum angle of incidence of θ^{\dagger} , the average deviation in retardation is computed as follows:

$$(\Delta V)_{\text{SIV}} = (k/\pi \theta^{\frac{1}{2}}) \int_{0}^{\theta^{\frac{1}{2}}} r^{2} 2\pi r \, dr = \frac{1}{2} k \theta^{\frac{1}{2}}$$

$$k = \frac{V_{0}}{k n_{0}} \left(\frac{1}{n_{0}} - \frac{1}{n_{0}}\right) \qquad \text{(occreated plates)}$$

In other words, the everage $\Delta\delta$ is one-half the maximum $\Delta\delta$ for a circular

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come of rays. The solid angle on of this come is

Hence over the cone of divergent rays, $\sqrt[5]{}$ varies from its mean value to $\mathcal{E}_{0} \triangleq \langle \Delta^{5} \rangle_{\text{eV}}$. The solid angle permitted, and hence the light gathering power depends on the permitted variation in $\Delta^{5} \approx \Delta p$.

$$\frac{a_{\text{N}} = \frac{a_{\text{N}} = \frac{a_{\text{N}}}{n_{\text{O}}} = \frac{(\Delta \delta)_{\text{mex}}}{\delta_{\text{O}}}}{\frac{a_{\text{N}}}{n_{\text{O}}} = \frac{a_{\text{N}}}{n_{\text{O}}} = \frac{a_{\text{N}}}{\delta_{\text{O}}}$$

$$\frac{a_{\text{N}} = \frac{a_{\text{N}}}{n_{\text{O}}} = \frac{a_{\text{N}}}{\delta_{\text{O}}} = \frac{a_{\text{N$$

Since the resolution R will be $\zeta_0/\Delta \zeta$, the general result for the luminosity will be

$$L = \frac{T_p A_p}{R} \frac{4\pi n_0^2 n_e}{n_e - n_0}$$

With "p = 0.17 the ratio of the luminosities of polarization and Fabry-Perot interferometers is approximately

$$\frac{0.17}{5} \left\{ \frac{4\pi \ n_0^2 \ n_0}{n_0 - n_0} \right\}$$

35 for AIP

10 for calcite

175 for quarts

The luminosity of the polarisation interferometer is thus greater than that of the Fabry-Perot, which is in turn more luminous than the grating

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or prime. It was pointed out in the detailed discussion of the last beobaical report that the disadvantage of the polarisation method was that for high accuracy it was restricted to two perfectly monochromatic lines. The higher luminosity may serve to rescue this instrument for practical use where luminosity is a heavy requirement, where the lines are quite narrow, and where the highest accuracy and contrast are not required.

10. SOURCE AND SAMPLING METHODS FOR DESTRETUM ANALYSIS

In order to use the spectrochemical analysis method on hydrogendeuterium mixtures, the sample may be in the form of either hydrogen gas or water. Bither the gas or the water vapor may be excited in an electrodeless discharge. The electrodeless discharge in a glass tube is preferred because of the possibility of contamination of metal electrodes. Someiderable time might otherwise be lost in cleaning out the discharge tube between samples.

Since only a few mm of vapor or gas in a small discharge vessel. As required, only a minute amount of sample is required. This is one of the advantages of spectrochemical methods over other methods of applicate.

Apparatus for bandling either type of sample have been constructed in our laboratories. The final choice probably depends most strongly on the actual form in which the sample for analysis is directly available. For the medical tracer applications the sample in usually liquid water. One or two drops suffice for the determination. The apparatus for the water vapor discharge is somewhat simpler than that for the gaseous discharge, but this is only if provision for dilution of semples for calibration, or otherwise, is required.

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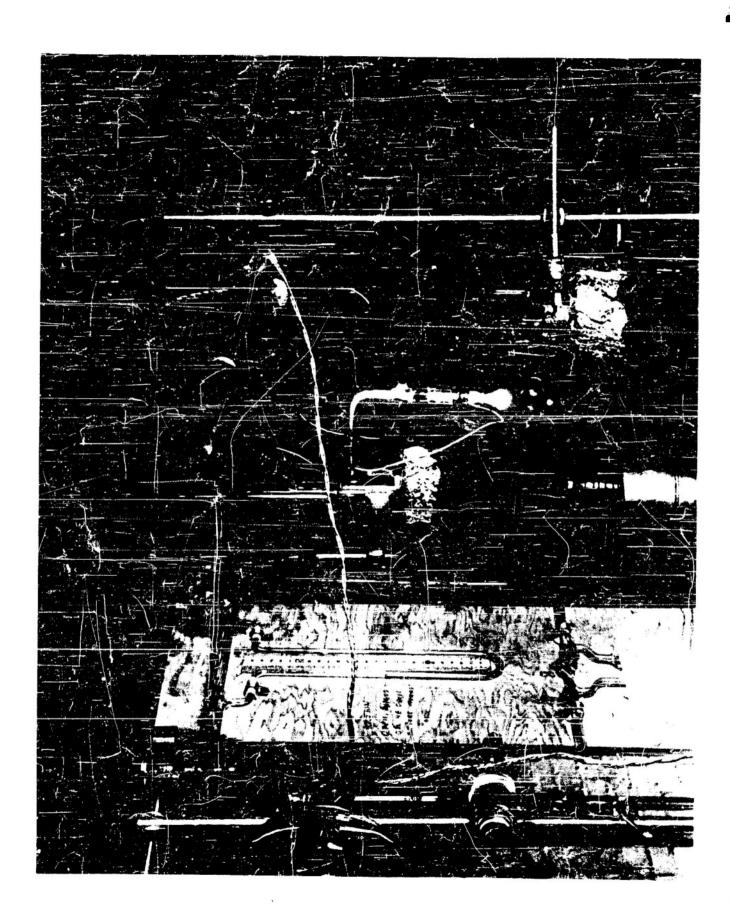
The apparatus for mixing and discharging hydrogen and derterium gases has been described in detail by Broids and Heyer of the Estimal Sureau of Standards (J. Opt. Soc. Am. 12, 37 (1932), who made an extensive investigation of the intensity ratio with pressure and with exciter intensity. Our apparatus was somewhat similar to their design, except that only one palledium tube was available. The photograph of our apparatus is shown in Figure 22.

The convenience of the water vapor method induced us to do most of the testing of the interferementer with this source. The apparatus is shown in Figure 23. A 125 watt microwave magnetron generator at 10 cm wave length was used for the excitation of the discharge.

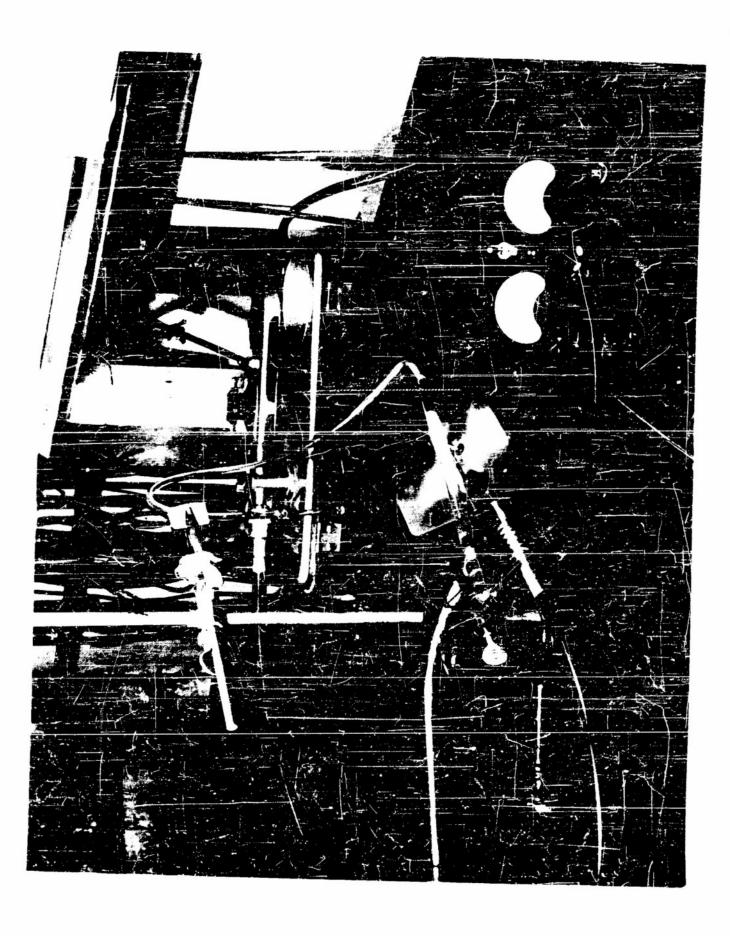
11. DEROVEMENT OF COSTRAST

If one attempts to measure devictive down to the low concentration available in ordinary hydrogen, a rather extreme contrast of 6000 to one is required. By careful optimisation of the ordinary techniques of Fabry-Perct interferometry it has been just barely possible to achieve this contrast. Since this problem may occur in even greater severity in similar isotope measuring problems, on other significations, it is another to inquire whether there are available any techniques which will push the contrast to even higher limits. Only two such methods have so far been uncerthed, and these were suggested to the author by Professor Jacquinot.

1. Self-Absorption. If light, after being writted, passes through a long path of the same gas, it may be absorbed. In the case of the strong and weak line, where high contrast is needed, the self-absorption will be much greater for the strong line. Therefore after traversing the path,







the two lines will be more nearly equal in intensity, and there will be an apparent decrease in the ratio of the light intensities. The situation in K-alpha is represent more difficult because the final state is not the ground state, therefore the line is not considered to be one susceptible to a high degree of self absorption. That is, before the absorbing path can absorb, there must be present a high concentration of hydrogen atoms in the first excited state. This means that a vessel of hydrogen gas alone will not serve as the absorber, but that the absorbing gas will have to be electrically excited as well as the source. Jacquinot has constructed long discharge tubes with sultinua reflections in which the ratio of D-sipha light to M-alpha was considerably increased. The excent of self-absorption, of course, will have to be accurately controlled if one is to be able to arrive at a measure of the concentration ratio from the intensity ratio. One of the beauties of the spectrochemical method of M-D analysis was that the intensity matic was almost exactly the same as the concentration ratio, so that extensive calibration was not required.

2. Interferometers in Series. If the required contrast eaunot be obtained with a single interferometer, the use of two interferometers in series should give an overall contrast equal to the product of those obtained with the single instruments. The difficulty have is in maintaining the tuning of the two instruments so that they are simultaneously on the peaks of the fringes. The stales spacing of course has to be maintained to a fraction of a wavelength of light. The air pressure tuning method was very useful in the single interferometer case. Probably in order to use two interferometers, they should be installed in separate vacuum tanks with independent control.

Application of this double monochromator principle is not of course limited to two identical instruments in series. A Fabry-Peret could be used in ceries with a prism, grating or other dispersing instrument of sufficient resolution. One of the best suggestions along this line appears to be a fore-dispersing element consisting of a solid mice-spaced Fabry-Perct plate, permanently tuned to the D-wavelength. After passing through this device, the intensity ratio presented to the Fabry-Perct would be much reduced, and such high congrast would not be required. Quests is a possibility for this solid spacer fore-dispersor. About 0.015 inches thickness is required (air spacer thickness divided by the index) to separate H and D by exactly 1/2 order. The advantage of quartz is its insensitivity to temperature, humidity or chemical attack. However, lithium or calcium fluoride is probably more desirable because of lower index, homes greater thickness, and because they are available in large synthetic pure crystals at low cost, and are isotropic media. The extresely thin plates required raise problems in standard finishing and polishing methods. Mica would be easier to prepare, but sight suffer from considerable absorption in such thicknesses,

Another possibility for the fore-dispersing element is one section of the polarization interference filter described in the preceding report. This could be of ADY, calcite or quartz.

If an adequate fore-dispersing element of the type suggested above can be manufactured simply and cheeply, an immediate suggestion is to dispense with the air spaced pressure tuned interferometer and use two such devices in series. The reference beam for the measurement could be

taken from the undispersed source, with a simple narrow band multilayer interference filter used to select the H and D light. After calibration with known samples, a direct reading method would be provided without necessity for scanning, nor for adjustment of parallelism of the plates.

The development of an adequate technique for manufacturing solid spacer interferemeters to the required tolerances in mice, quarts or the fluorides suggests an important proposal for future experimental work in this field of high resolution interference spectroscopy.

12. FROTCALS FOR FUTURE WORK

The report on the hydrogen-douterism enalysis problem by means of the Pakry-Perot interference has demonstrated that this instrument has the possibility of the highest luminosity, resolution and contrast for the spectral analytical determination of isotopes. The hydrogs - deuterium isotope shift is of course one of the largest. It is proposed that future work he aimed directly at separating other isotopes of much narrower isotope shifts, notably uranium, because of its importance.

With the heavier elements, although the isotope shift is smaller, the Doppler breadening would be less. Therefore one of the major limiting factors will be the flatness of the interferometer plates. Considerably more work needs to be done both on the techniques of preparing flat quarts plates and on the techniques for testing them. The testing is necessary to the preparation, since the precise optical worker can make fine corrections if he is given very precise information as to where the defects lie. Baird Associates is fortunete in being able to obtain the services of Mr. Wagn Hargbol for this works. The problems have already been discussed,

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and Mr. Margbol is anxious to acquire the best apparatus and techniques for achieving the ultimate in such plates.

Further work meets to be done on the improvement of the techniques of laying down multilayers. This involves purifying the disloctrics
evaporated, extensive gudgetry for rotating plates in the evaporator to
insure uniformity, solution of the exact theoretical equations for the
variation of transmission with thickness of deposited layer so that the
optimum monitoring system may be devised, etc.

Another major problem requiring extensive research work is the problem of sources for the heavier elements, such as uranium. The vapor pressure is low, and it is difficult to excite them, even in arcs or sparks. In addition, such arcs and sparks are notoriously sources of extreme fluctuations in light intensity compared to gaseous discharges. Hence we may expect the problem of noise to be intensified with these elements.

An additional complication may be introduced because of the complexity of uranium spectra. For N-D, a 50 angetrom vide multilayer interference filter is quite adequate as an isolating filter. For uranium, it may be necessary to do research on the development of more narrow isolating filters.

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APPRENTY T

THE DETERMINATION OF EMAYY WATER BY IMPRANSO ABSORPTION INCRMIQUES By W. A. Patterson

SUPPLANT

Using 0.5 mm Calcium Fluoride cells, the differential method of infrared analysis, a reversed cell technique and wide fixed slits, Deuterium Oride as an additive in water can be readily detected in percentages as low as 0.01% with a probable precision of + 0.001%. Under present conditions an ultimate sensitivity of 0.00% Deuterium Oride seems probable. The determination of ultimate sensitivity has been restricted by bad scattered light conditions when spectrophotometric slits are widened. If this scattered light problem could be solved, then an ultimate sensitivity of 0.001% Deuterium Oride does not seem out of the way.

Introduction

Sensitive methods for the determination of heavy veter (Deuterium Oxide) in the presence of ordinary water have been based either on the mass spectrometer or some light dispersing instrument such as an existing spectrograph or monochromator using electrodeless gas discharges. Thomas 15 describes a mass spectrometric method for PgO in EgO which appears to be accurate to \pm 0.2% in the range 0.2% to 1%, and \pm 0.3% in the range 1% to 10%. This method is based on the determination of the relative contentrations of mass 19 from DOM and mass 18 from EgO. Thomas indicates that the \pm 0.2% limit of accuracy might be due to adsorption of heavy water on the vacuum system walls. This had been realized by Orchin, Mender and Friedel¹⁶

^{15.} B. W. Thomas, Anal. Ches. 22, 1476 (1950).

^{16. 0.} Orchin, I. Wender and R. A. Friedel, Amal. Cham. 21, 1072 (1949).

Peira Associates, Inc.

whe, in order to get increased accuracy, converted solutions of D₂O and H₂O to Methyl Deuteride (SH₂D) and Methane (SH₂) before using the mass spectrometer. This method gave values for beevy water to within O.15 in the range 1.4 to 15%, but neglecting the errors inherent in the "conversion", the reproducibility of the mass spectrometer values on the Methyl Deuteride was at least a factor of 2 better than this.

The Orchin et al method is interesting in that it was applied to mixtures of hydrocarbons of relatively high molecular weight, some compounds of which had Douterium atoms substituted for Hydrogen atoms. These hydrocarbons were accurately converted to water and Deuterium Oxide by a method developed by Keston, Rittenberg and Schoenheimer¹⁷. This technique has important applications and means that a good method for determining D₀O in H₀O could have wide spread applications.

The papers of Broids¹⁸, ¹⁹ appear to be the most authoritative using emission techniques. In the first paper¹⁸, gaseous mixtures of Bydrogen and Deuterium from 85% to 100% Deuterium were analysed to within 0.1% of the major component. In the second paper¹⁹, with the same instrumentation, and applying isotope dilution techniques the total water content of materials such as "hydrates" was determined to within 3% of the amount present. This paper illustrates at least one other type of application for mathods determining D₂O in H₂O. According to the authors, the technique is perfectly general, and applications were planned for the study of water binding and transport in biological systems.

^{17.} A.S. Keston, D. Rittenberg, and R.J. Schoenheimer, J. Rio. Chem. 192, 227 (1957).

^{18.} M. P. Broids and J. W. Moyer, J. Opt. Soc. Am. 12, 57 (1932).

^{19.} E. J. Morowitz and H. P. Broide, Anal. Chem. 24, 1657 (1932)

solves to the use of very small samples, for example as little as 5 milligrams. This does not mean, however, that these instruments are the most sensitive for the detection of lowest concentrations of Deuterium Oxide in water. On the mass spectrometer or the emission spectrograph, for highest sensitivity, it is advisable when handling very small samples of water for analysis, to reduce the aqueous material to the Deuterium and Mydrogen gases. Dubbs and Shinard and Erms 21 have described methods for this purpose, using but Zinc as the reducing agent. These papers would definitely have a bearing on the applications of any emission type instrument we might develop.

The mass spectrometric methods have disadvantages in that the instrumentation is expensive and for the most sensitive results the sample needs suitable proparatory procedures such as reduction to the elemental games or the coversion to Nothyl Deuteride and Nothane. Spectrographic type instruments are cheaper but high sensitivity on low exacentrations of Dauterium Cride in water has not been proven as yet. Eare again conversion to the elemental games seems desirable. Spectrographic equipment is also not widely distributed emeng those chemists who are most likely to require Deuterium Cride analysis.

More sensitive methods seem desirable, methods which could either be applied to instruments more widely distributed among organic

^{20.} C. A. Dubbs, Anal. Them. 25, 828 (1955).

^{21.} F. P. Chinard and T. Ems, Anal. Chem. 25, 1415 (1955).

According to further information received from Dr. J. L. Saunderson, this has been done by Broids and company at the Mational Bureau of Standards, but is unpublished to date.

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chemists, and biochemists, or the principles of which might hand themselves to the development of simpler and less costly instruments. Improved exission methods have been under consideration in this laboratory, and research undertaken. There remains a further possibility, and that is infrared absorption which heretofore has not been given serious consideration, either by us or by others.

In 1950 Thornton and Condon described an infrared method of low sensitivity for the range 0 to 100% Dauterium Oxide in water. An accurrecy of 3% was claimed, but it was expected that this could be improved upon. The method was based on the shift of the 2.5 meron water band to 3.8 microns in Deuterium Oxide. Treamer and Walker 20 made a more thorough investigation of the Thornton and Sondon method and discovered that the intensity of the 3.8 microns band for Deuterium Oxide was very sensitive to temperature changes. They developed a thermostated cell with which it was possible to analyze for Deuterium Oxide, with an accuracy of + 0.15 on 3 Desterium. The Tremer paper is slightly vague in that the authors speak constantly in terms of Douterium, though the analysis is based on heavy water. For heavy water, the above values of Douterium should be multiplied by nine. This work was done on a single beam instrument and the analytical range considered was 0 to 5% Deuterium (0 to 45% Deuterium Oxida). These authods were not sensitive and the full potential sansitivity of the infrared technique such as might be obtained from a double bear instrument and differential methods had not been studied. These seemed worth investigating.

^{22.} V. Thornton and E. I. Sondon, Anal. Chem. 22, 690 (1950).

^{25.} N. R. Tremer and R. W. Walber, Instrument News, 1, No. 1, 1 (1992).

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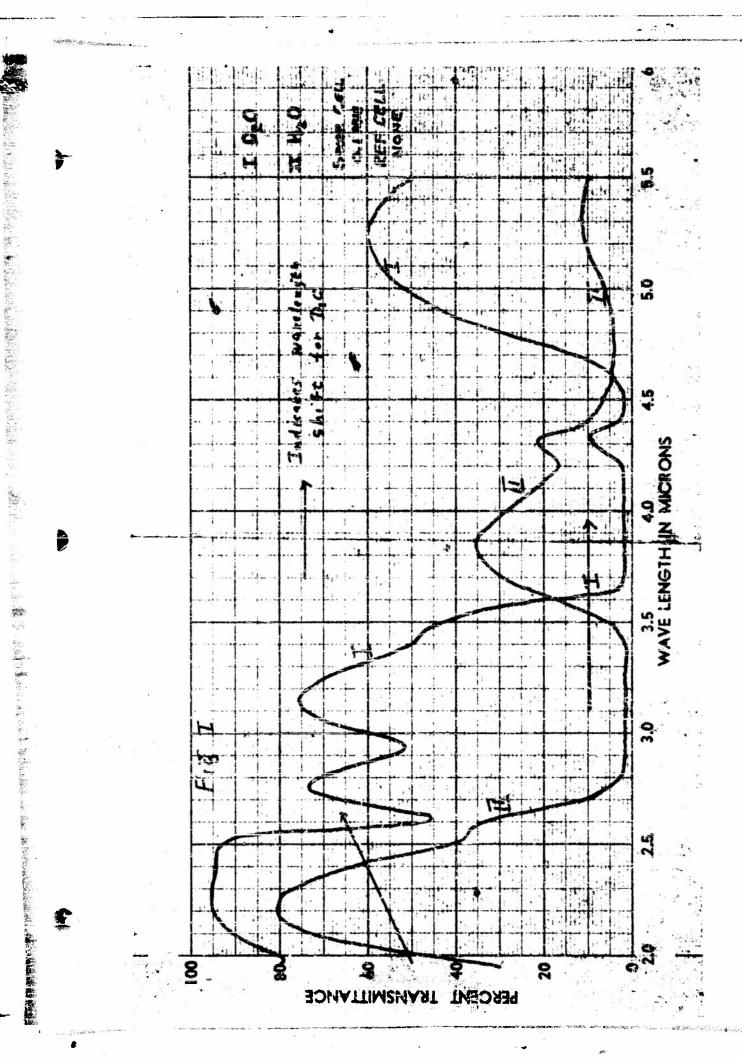
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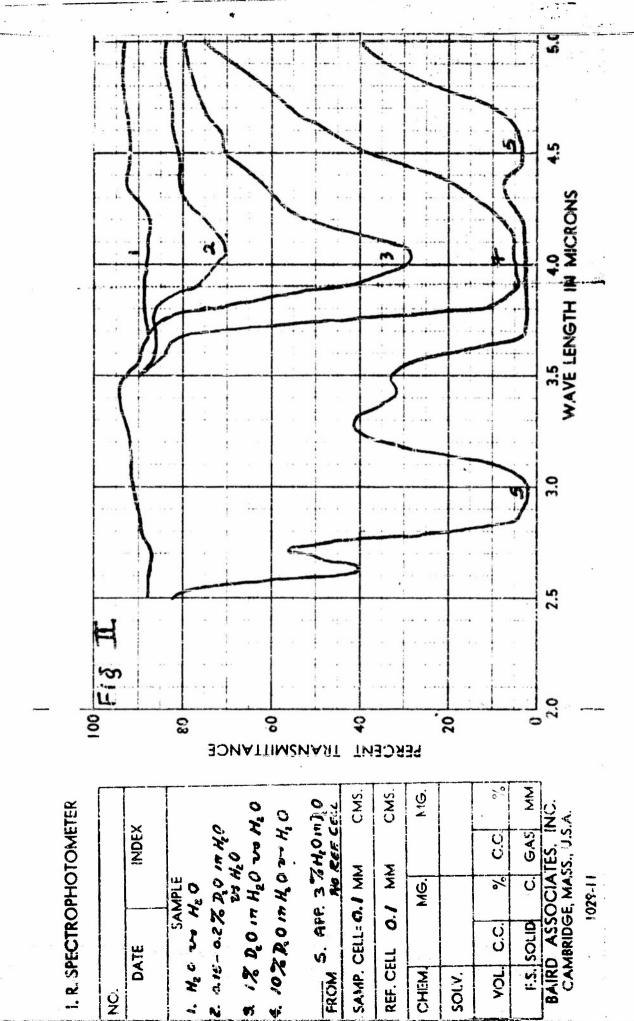
The object of this investigation was to apply the most emplicies techniques of infrared spectroscopy to reach some conclusion on what are the lower limits of detection of Deuterium Oxide in the presence of water. In this initial work, precision and reproducibility were not considered. These would necessarily be the subject of another and more extensive investigation.

Experimental

1. The Use of O.1 mm Salls

The most suitable thickness of an infrared absorption call for water is scrething less than 0.01 mm depending on the problem. Transcr was able to use a thickness of 0.02 mm for the measurement of Deutsrium Oxide. Any thicknesses greater than these would give spectra showing very high absorptions in most regions from 2 to 16 microns. Figure I shows the spectra of water and heavy water in 0.1 mm Galdium Fluxride calls with a Salcium Fluoride prism. The use of Calcium Fluoride cells is necessary bycause of the solubility of rock selt in water. However the use of the Salcium Fluoride prism in this particular instance had no significance. Studying the spectra, it is evident that the strongest checrytion bands for water and Douberive Oxide are at about 3.0 and 4.0 microns. The 4 micron Diuterium (mide band appears at a gap in the absorption of the water spectrum and would be suitable as an analytical band which is in agreement with the choice of other workers on this enelysis. The 2.9 micron band in the heavy water spectrum is evidently due to something less than 0.5% water as the Deuterium Oxide was only guaranteed to be at least 99.5%





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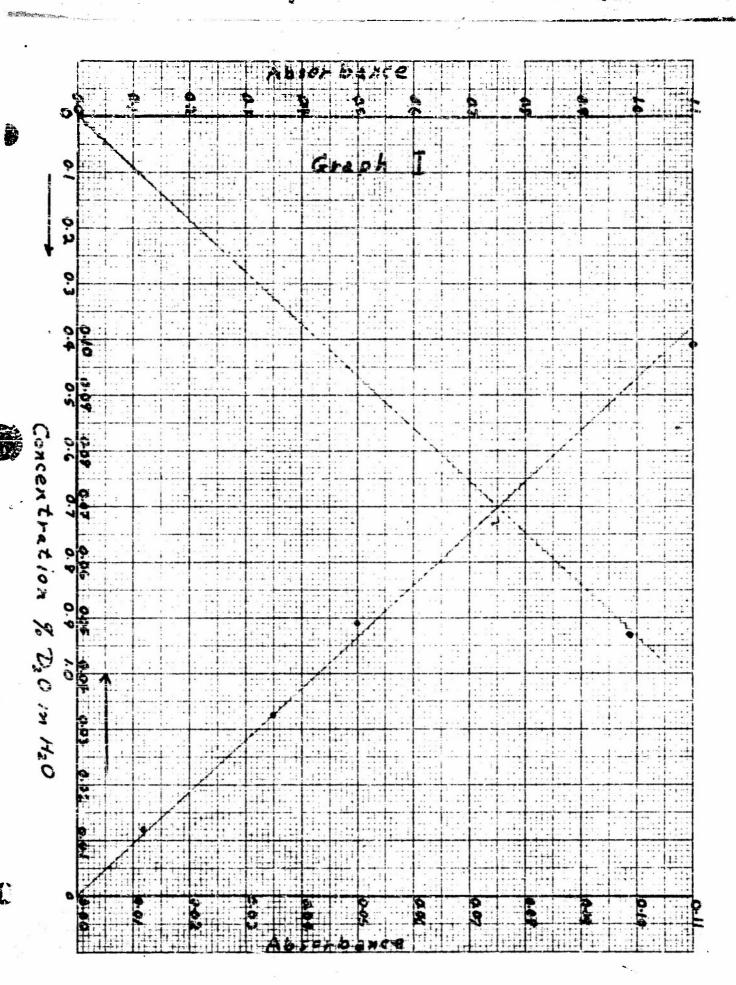
On Figure IV are shown the six standards run against distilled water in the reference beam, using matched 0.2 mm cells. These spectra were made using fixed slits, the width of which had been adjusted to give a D.S. voltage of 2.0 volts at the mechanical rectifier, at a wavelength setting of 3.7 microns, and with an 0.2 mm cell containing water in the sample beam and an opaque in the reference beam. Gains of voltage and known amplifier were as usual. This was also the procedure in edjusting the slit widths in subsequent work, although the actual voltage figure will very.

Using as the base lime the curve for H water against water, the absorbance values of these absorption bands were calculated. The results are plotted on Graph I. Observe the nice straight line relationship and the high degree of correlation between the different results; even when the 0.01 to 0.1% D₂0 in H₂0 range was expanded as shown in the right hand curve.

3. The Use of 0.5 am Celle

Since 0.5 mm cells instead of 0.2 mm cells should in theory give increased sensitivity, this was the logical next step. The spectrum of water in an 0.5 mm cell showed 99% absorption in the 5.8 to 4.0 micron region. This did not look very hopsful. Proceeding however to open up the slits to 0.5 velts as per our previous mentioned adjustment procedure, attempts were made to run water against water in the 0.5 mm cells. The very wide slits, however, introduced a complication in that much scattered light was introduced. This caused considerable trouble. For example, between 5.0 and 5.4 microns where there should be complete absorption in

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both bears, there was a strong does scale drift, which could be reversed as the slite were narrowed, indicating definitely that there was an unbalance due to scattered light when the slite were set for the 0.5 volt reading. The instrument had a normal up scale drift. The does scale drift was so strong that it played bears with all the survey and note was reproducible. It was not even possible to get a reproducible water we, water baseline or reference to determine whether the two cells were seatched, though of course tills could be checked with other solvents. It should be noted that this dress scale drift only occurred when there was complete absorption. Where absorption was not complete the pass them want up scale, trying to go back to its proper balance position. This meant that there was little chance for recording at wavelengths prior to the actual absorption band, but had to record immediately into the band or where there was energy, which in this case amounted to the same thing.

Considerable time was spent in studying this scattered light problem, full details of which were given in notebook 27400. A description of the tests carried out would only complicate this report, and besides it is an instrumental problem rather than an analytical one. Though it is the key to getting the highest sensitivity. All that need be said at this time is that some partial success was obtained even with slit settings of 0.5 volts at 5.7 microns. Each was still to be desired however, and it is quite probable that the sensitivity attained in this work could be at least doubled if a completely setisfactory ensure to the scattered light problem were to be found.

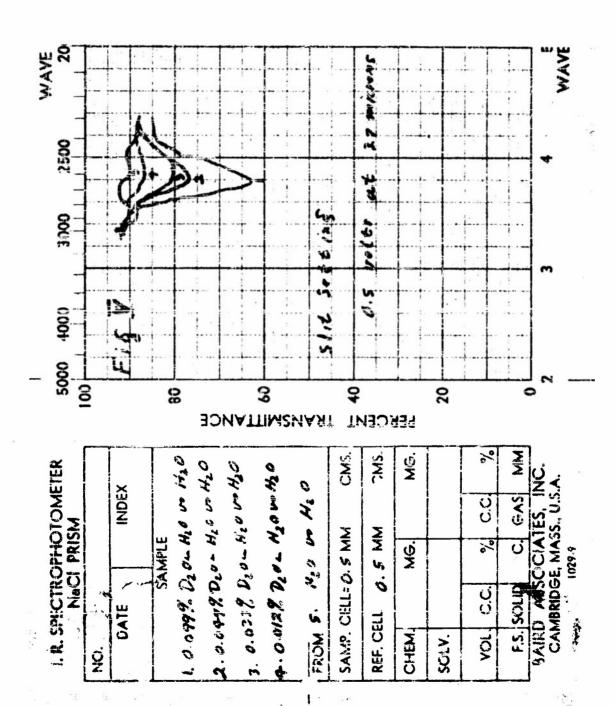
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Pigure V shows the first successful curve, rea differentially, with Dauterium Oxide standards in the 0.5 mm calls. The absorbance values as shown in Graph 2 are only about 50% higher than those obtained using 0.2 mm calls. This increases in absorbance should have been about 2 1/2 times. The low values can be attributed almost satisfully to the aluggishmess of the pen due to the low energy levels. It should be realised that the 0.5 volt setting was at 3.7 microns which meant that at about 5.9 to 4.0 microns the smount of energy would be on the order of 0.1 volt or less in both busses.

4. Using Reversed 0.5 mm Colle

Robinson 21 has shown that increased sensitivity can be obtained by first running samples differentially in the normal way said them recording again above the first spectrum but with the samples reversed in the two bears. This gives two curves which are mirror images of each other, the distance between the two peaks being an amplified absorbance value for the concentration in question. To do this properly only the samples are reversed, and not the cells, and the test them ampenestes for thickness discrepancies between the two cells. In the present work this refinement of technique was not deemed necessary and both cells and solutions were reversed.

The successful emplication of the Boverest Cell technique in this particular problem (only) required a reliable 100% line as starting point for recording both the upward and document curves. This does not meen a true 100% line, but simply a starting point at which the pan could 24. D. Z. Robinson, Anal. Then. 25, 519 (1972).





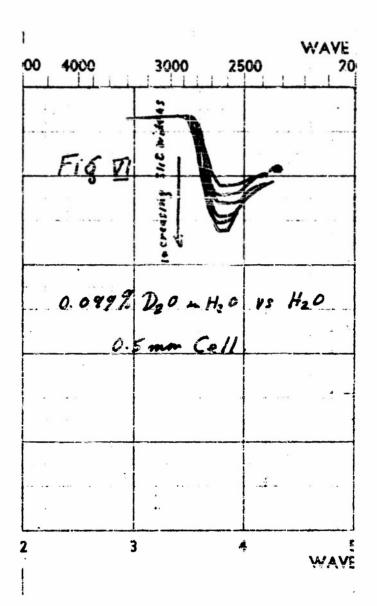
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be set and the position of which would not change during the recording until the scapping entered the D.O electrica band region of the spectrus. This mount no down scale drift from contiered light. It also meant a low noise level. To attein the latter the voltage amplifier was switched to the St position and to obtain a responsive pen without southered light interference the slite were gradually opened and curves recorded, noting the stability of the starting point. The series of ourves obtained, using the 0.09% Pentorium fundo standard, as the alite were videned are shown in Figure VI. So attempt was made to record the D.S. voltage on these alit openings. Using the alit opening designated by the lowest ourse on Figure VI, four standards were recorded along with water with the reverse cell technique using 0.5 mm cells. The results are shown in Figure VII. The flat maxima of these curves is the to lack of sufficient energy betrem the two beams activiting the serve mechanism. Figure VIII show the curve cotained from the same standard using A.S. bias in which the curves take their normal shape.

5. Discussion

On Graph 2 are plotted the absorbance values for the three standards 0.012, 0.055 and 0.05% D₂0 in H₂0 which were obtained during the course of this investigation as increasingly consistive methods were applied. These absorbance values are from Figures 27, V and VIII. As mentioned previously the expected increase of 2 1/2 times the absorbance values on going from 0.2 to 0.5 am thick calls did not occur. As also mentioned previously this is very probably due to the low energy levels involved which had to be kept low as wide alit openings introduced scattered light problems.



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the results are completely in line, being about twice there obtained for the standard method using 0.5 mm calls (curve 2). Note that curve 3 data not go to sure but to 0.015 absorbance value which is the absorbance difference between the two calls filled with water as indicated in Figure VIII. If solutions had been reversed but not the calls, this curve 5 of Graph 2 should have gone to sero, as with curves 1 and 2. A glance at the about contense bands of Figure VIII and at curve 5, Graph 2 makes it readily apparent that 0.015 Deutswices Oxide added to ordinary water is readily measurable. It is also apparent that concentrations as low as 0.00% Deutswices on curve 5, Graph 2 indicate that the percent relation between the points on curve 5, Graph 2 indicate that the percent deutswices Oxide in the range 0.01 to 0.0% could probably be determined with an accuracy of \$ 0.0015 if a precisely controlled method were Sevelepod.

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While final conclusions abould not be drawn from a single series of experiments, there seems little dealt that the infrared absorption method is fully expelle of detecting and measuring 0.03% Desterium Coids in water. The possible degree of precision in the analyses could be about 4 0.003% or batter.

Work so far has indicated that percentages of Deuterium Oxide as low as 0.00% are within the sensitivity limit of this method. The ultimate sensitivity of the infrared method has not been resched in this work due to restrictions on the limits of the experimental technique from scattered light. If these were removed a sensitivity of 0.000% does not seem out of the way.

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APPENDIX II

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- 2.1 2.8 Previous Technical Reports Submitted by Baird Associates, Inc.
- 5.1 5.16 Published Papers and Talks. (See detailed list and abstructs, Section I of this report.)

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